Nuclear incompressibility from the viewpoint of spherical and deformed QRPA calculations







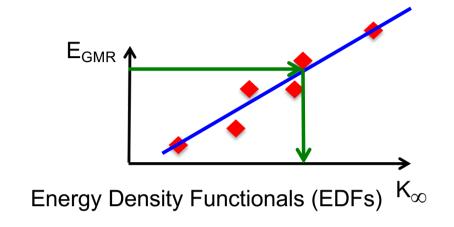






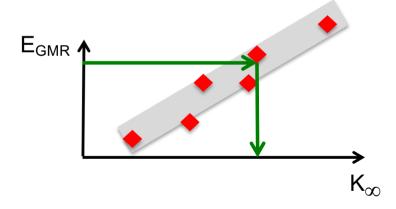


Background



Only self-consistent calculations that treat uniform matter and the response of finite nuclei on equal footing allow extracting K_{∞}

J.P. Blaizot, Phys. Rep. 64, 171 (1980)



There are different sources of model dependence in this procedure.

One **key point** is that different EDFs have different assumptions for the density dependence.

GC *et al.*, Phys. Rev. C70 (2004) 024307.





- Sensitivity to the choice of the nucleus?
- Can we understand what happens if we apply this procedure to superfluid/deformed nuclei?

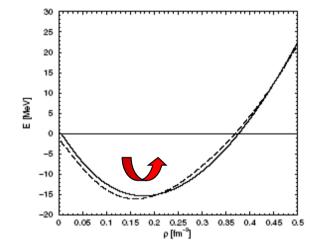
(Q)RPA using EDFs in a nutshell

$$E = \langle \Psi | \hat{H} | \Psi \rangle = \langle \Phi | \hat{H}_{eff} | \Phi \rangle = E[\hat{\rho}]$$

 $\ket{\Phi}$ Slater determinant $\Leftrightarrow \; \hat{
ho} \;$ 1-body density matrix

 $H_{eff} = T + V_{eff}$. If V_{eff} is well designed, the resulting g.s. (minimum) energy can fit experiment at best. Hartree-Fock or Kohn-Sham.

• Within a time-dependent theory (TDHF), one can describe harmonic oscillations around the minimum.



• The restoring force is:
$$v\equiv rac{\delta^2 E}{\delta
ho^2}$$
 .
$$X_{
m ph}|ph^{-1}
angle -Y_{
m ph}|hp^{-1}
angle$$

$$X_{\rm ph}|ph^{-1}\rangle - Y_{\rm ph}|hp^{-1}\rangle$$

The linearization of the equation of the motion leads to RPA1.

¹Random Phase Approximation.

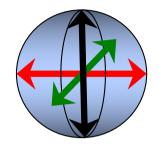
$$\begin{pmatrix} A & B \\ -B^* & -A^* \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = \hbar\omega \begin{pmatrix} X \\ Y \end{pmatrix}$$





Incompressibility of nuclear matter from ²⁰⁸Pb





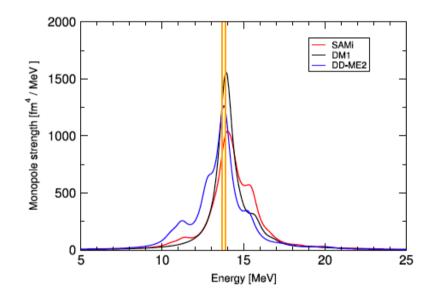
The compression-mode giant resonances and nuclear incompressibility

Umesh Garg a, Gianluca Colò b,c,*

Progress in Particle and Nuclear Physics 101 (2018) 55-95

240 ± 20 MeV from the GMR in ²⁰⁸Pb

EDF	SAMi	DM1	DD-ME1
K_{∞}	245	225	245.5
	MeV	MeV	MeV



Non-magic nuclei may (?) point to lower values of the nuclear incompressibility

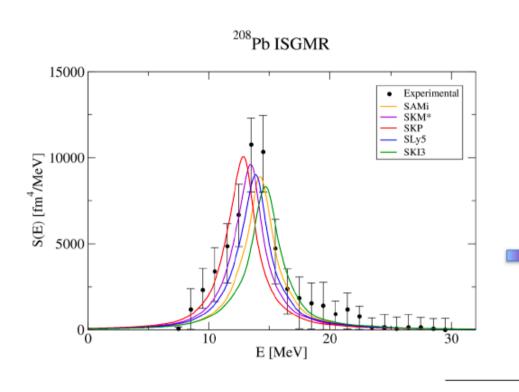




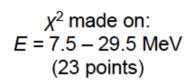
A different kind of analysis

by S. Zaghet





 $K_{\infty} = 220 - 230 \text{ MeV}$

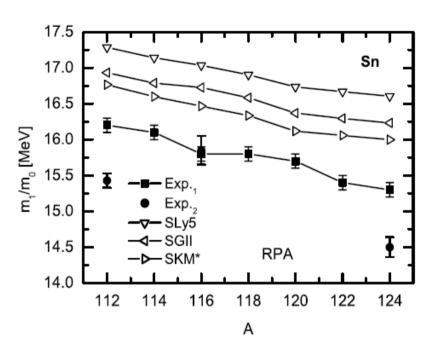


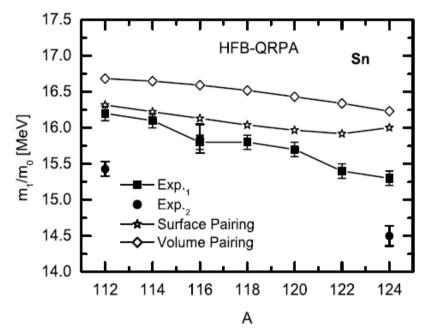
	SKP	SKM*	SLy5	SAMI	SKI3
χ^2	68.6	35.7	31.5	44.6	73.8
reduced χ^2					
K_{∞} [MeV]	201	217	230	245	258





Superfluid nuclei





Pairing correlations <u>lower</u> the energy of the GMR (right panel) with respect to RPA (left panel).

J. Li et al., Phys. Rev. C 78, 064304(2008); L. Cao et al., Phys. Rev. C 86, 054313 (2012).

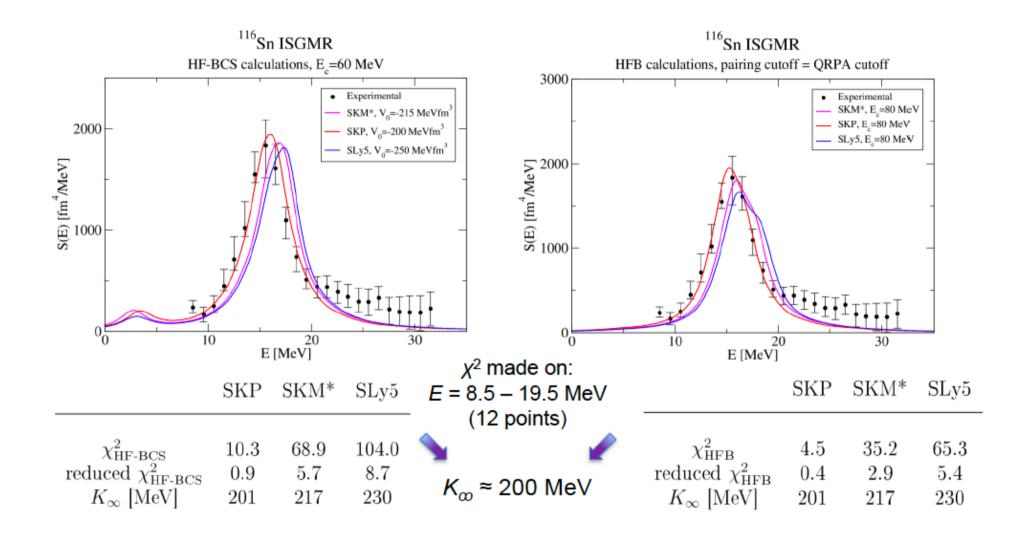
The effect on the incompressibility can be of the order of 10%.

E. Khan *et al.*, Phys. Rev. C 82, 024322 (2010).





- Quantitative results depend on the kind of pairing force
- The information from g.s. pairing gap does not fully constrain it

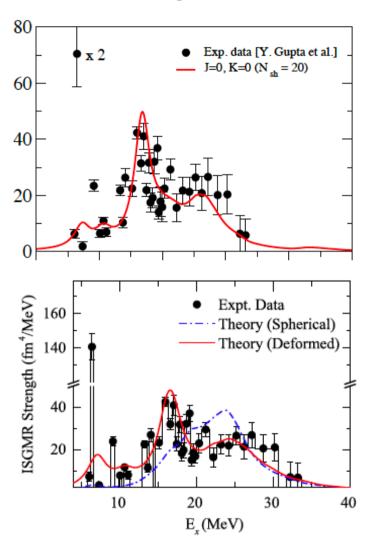






Well-deformed nuclei

²⁴Mg, SKM*



We compare with RCNP data from Y. Gupta *et al.*, PRC 93, 044324 (2016).

The two-peak structure is evident.

Thanks to K. Howard.

Similar calculations by K. Yoshida was used to show that the double peak is related to deformation.

General understanding: in deformed systems K^{π} are the only good quantum numbers. The K=0 components of the GMR and GQR are coupled and they both show up.





GMR Workshop (GDR RESANET, 11/2020)

Deformed QRPA implementations

Either HFB or HF-BCS equations with a Skyrme force and a pairing force are solved on a (HFBTHO / SKYAX).

> M. Stoitsov et al., Comp. Phys. Comm. 184 (2013) 1592; P.-G. Reinhard, SKYAX (unpublished).

This allows to determine study the potential energy surfaces (PESs).

$$E = E(\beta)$$

- The QRPA equations are solved on a discrete basis with good Kⁿ.
- Using two different QRPA approaches strengthens our conclusions.

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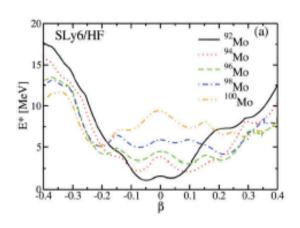


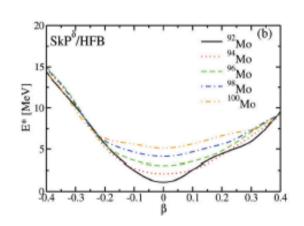


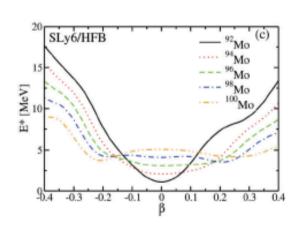
Isoscalar monopole and quadrupole modes in Mo isotopes: Microscopic analysis

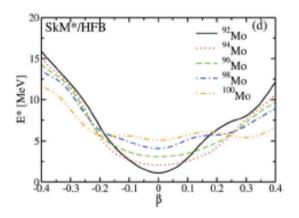


Potential energy surfaces of ^AMo









Pairing and deformation are linked to each other, we will not be able to disentangle the effect of either effect in the monopole strength.

With increasing A, the curves become more flat ("quadrupole softness").

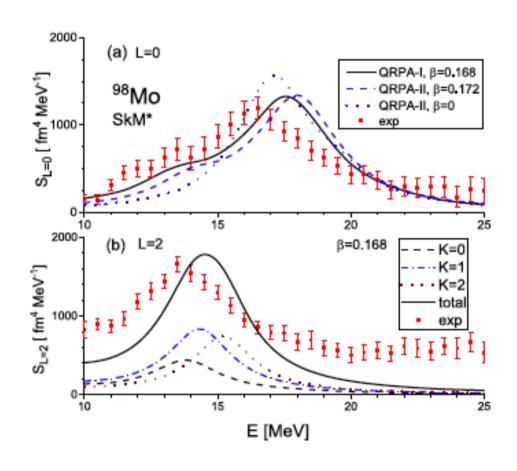
It is hard to obtain a straightforward indication of the g.s. deformation from experiment.

 $\beta = 0.109, 0.151, 0172, 0.168, 0.162 \quad (A = 92...100)$





Effect of (modest) deformation



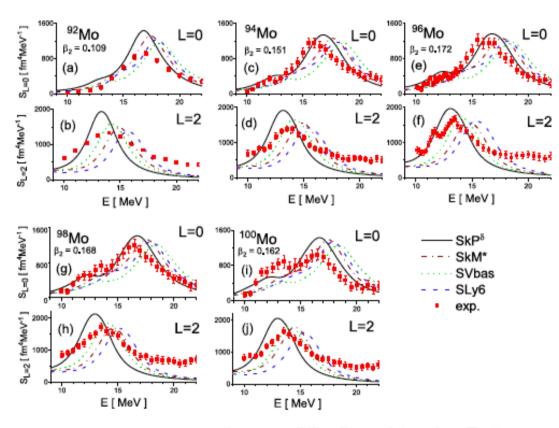
Implementing QRPA in two different manners we do not find results that change significantly.

Even a small deformation can shift the ISGMR by around 1 MeV.





Monopole and quadrupole strength in AMo



The "shoulder" is due to the monopole-quadrupole coupling.

The Skyrme EDF that better reproduces the GMR (GQR) results is SkP^{δ} (SVbas).

Warning, warning ...

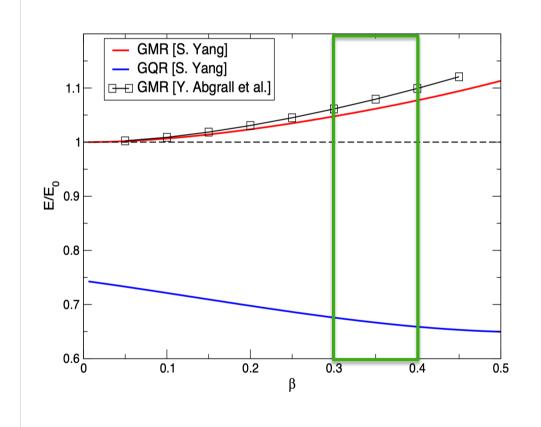
Incompressibility K_{∞} and isoscalar effective mass m^*/m for the Skyrme forces SVbas, Sly6, SkM*, and SkP⁸.

	SVbas	SLy6	SkM*	SMP
K _∞ [MeV]	234	230	217	202
m*/m	0,9	0.69	0.79	$\overline{}$





In <u>macroscopic models</u>, one starts from the "ideal" GMR and GQR of spherical nuclei and finds that two peaks arise from the **coupling of the** K=0 components of **GMR and GQR**.



• S. Yang, NPA 401 (1983) 303 Hydrodynamical calculations

$$E_0 \left(1 + 0.86\delta^2 - 1.25\delta^3 \right)$$

Y. Abgrall et al., NPA 436 (1980) 431
 Adiabatic cranking model

Quite similar results!

- T. Suzuki and D. Rowe, NPA 289 (1977) 461
- Y. Shimizu and K. Matsuyanagi, PTEP 72 (1984) 1017
- S. Åberg, NPA 473 (1987) 1
- D. Zawischa et al., NPA 311 (1978) 445





Conclusions

- Personal (or shared?) belief: the nuclear incompressibility should be around 240 MeV, if we base ourselves on few magic nuclei.
- In some cases (e.g. ⁹⁰Zr), experimental uncertainties play the largest role.
- Superfluid nuclei may point to a lower value but our knowledge of pairing is not accurate enough to say more.
- We may need to fit deformed nuclei into the picture.
- The coupling between the K=0 components of the monopole and quadrupole plays a role in determining the peak of the monopole strength.
- Perspective: isotopic/isotonic chains in which open questions could be benchmarked.





Thank you!



