

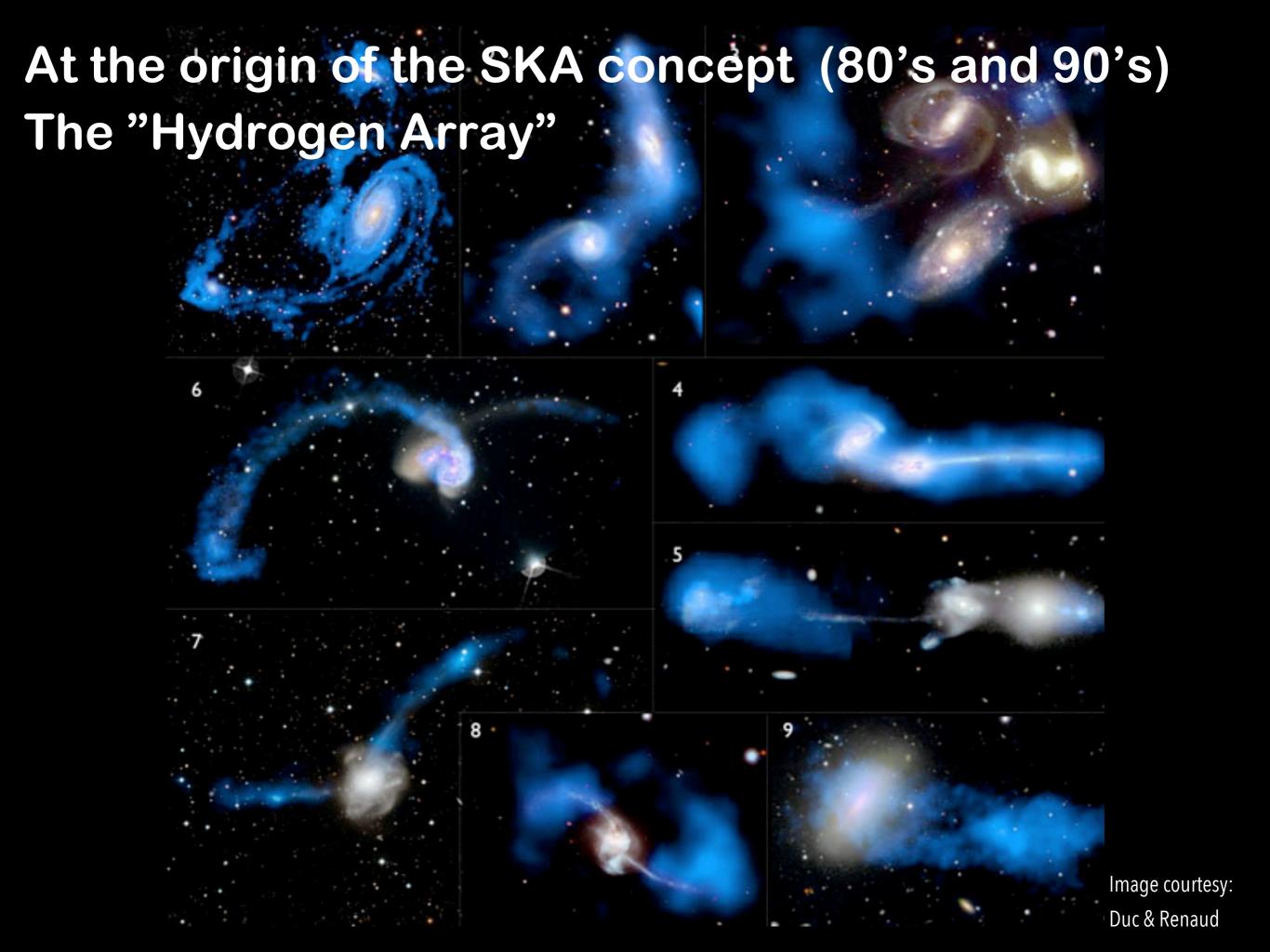
C. Ferrari

Astronomer

Director of Maison SKA-France







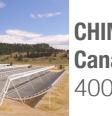
A Golden Age for Radioastronomy

SKAPathfinders

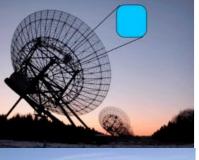


NenuFAR France 10-85 MHz

LOFAR Europe30-80 MHz +
110-240 MHz



CHIME Canada 400-800 MHz



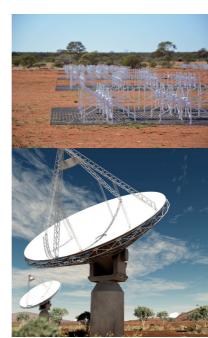
APERTIF The Netherlands1 – 1.750 GHz



EVLA USA 1- 50 GHz

And 10 more (in Europe, America, Japan, India, Australia)

SKA Precursors



MWA Australia 80 - 300 MHz



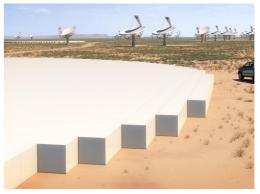


HERA South Africa 50 - 250 MHz

MeerKAT South Africa 0.580 – 14 GHz

SKA



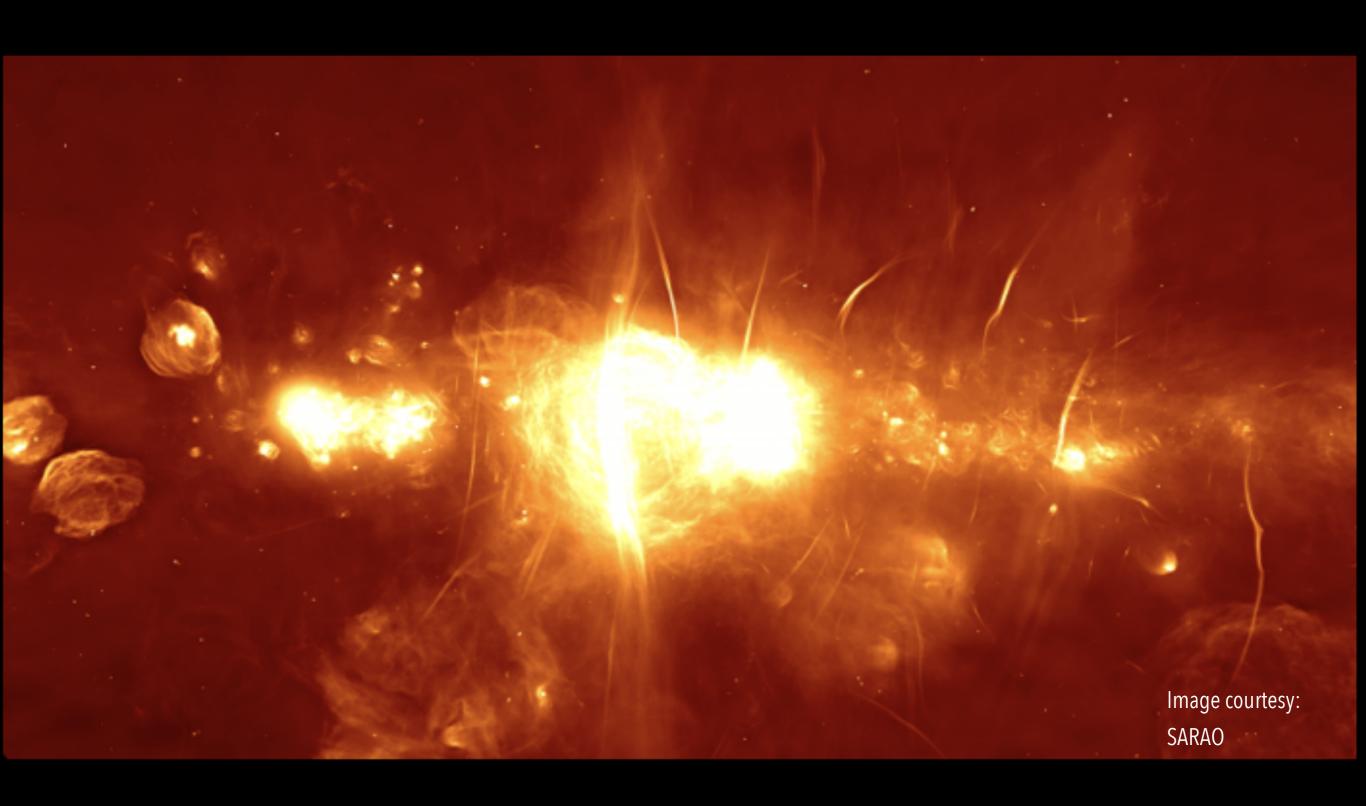


MID & LOW
South Africa &
Australia
50 MHz – 24 GHz

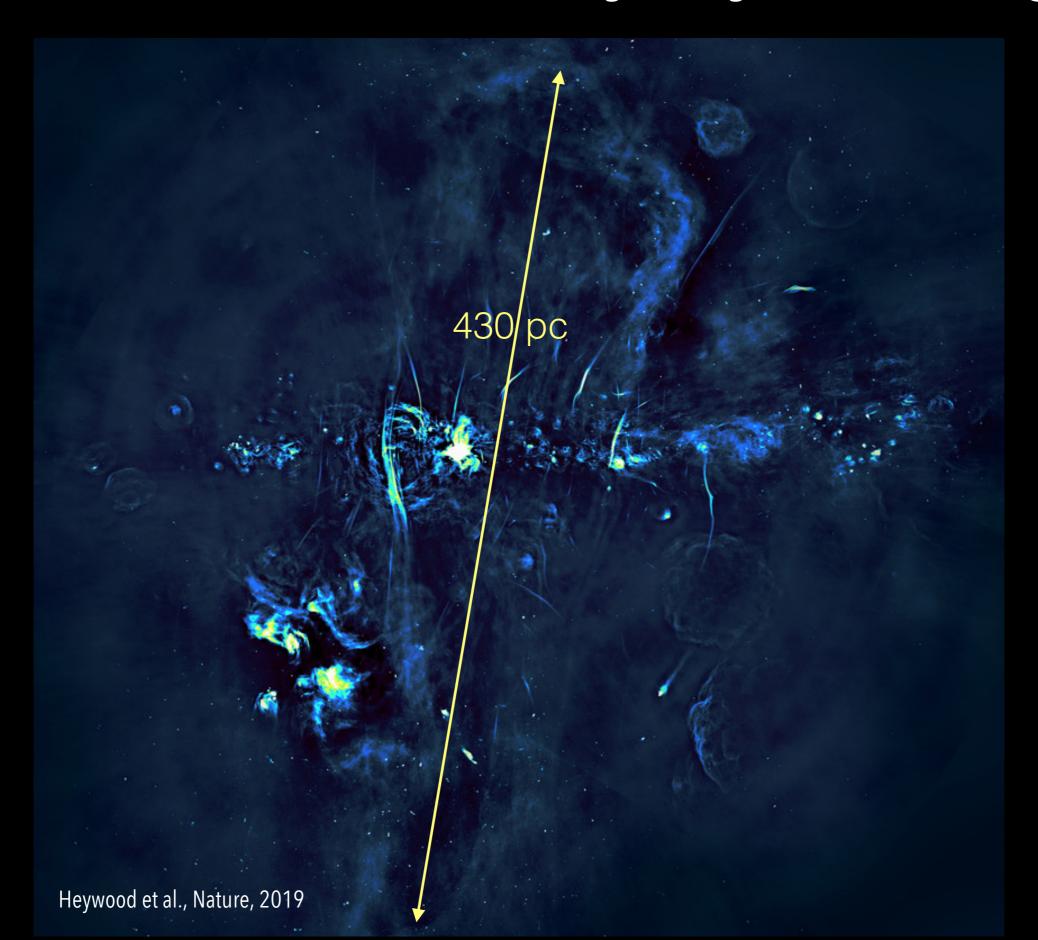




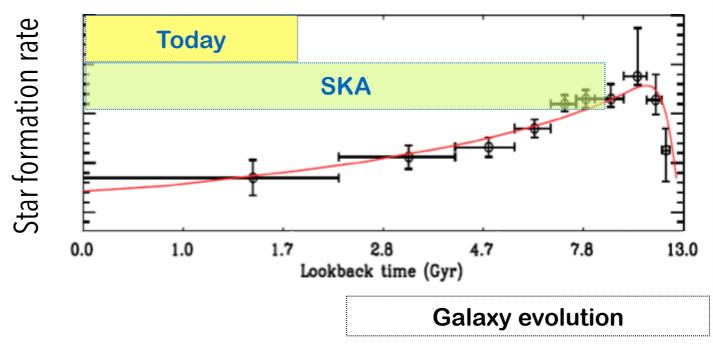
A MeerKAT view of the Milky Way central region

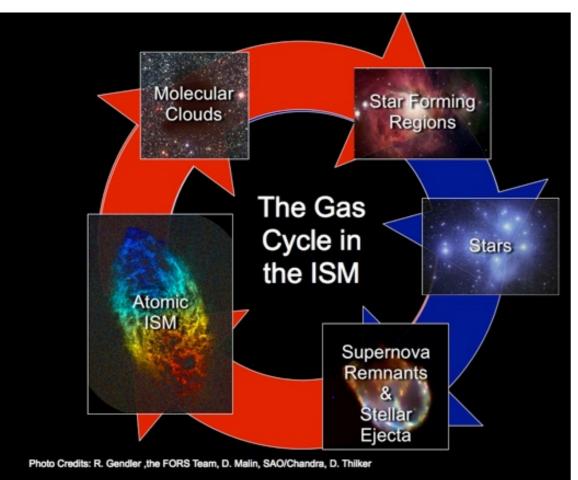


A MeerKAT view of the Milky Way central region

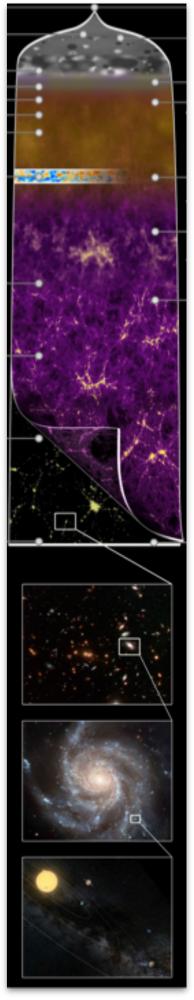


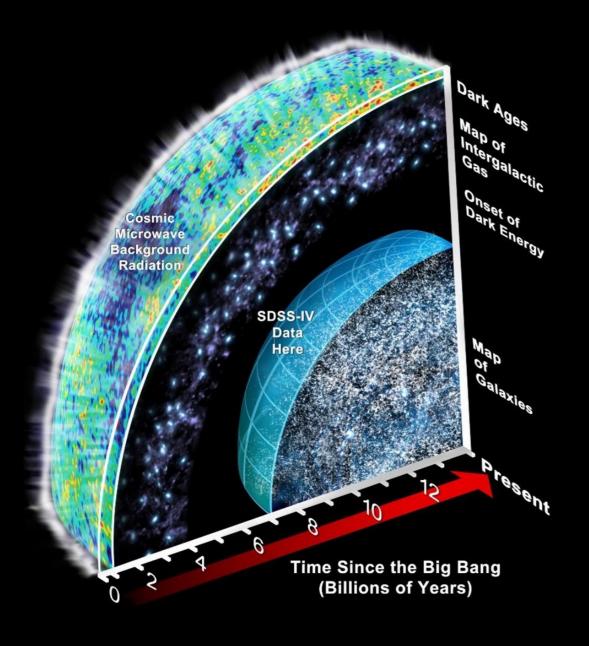


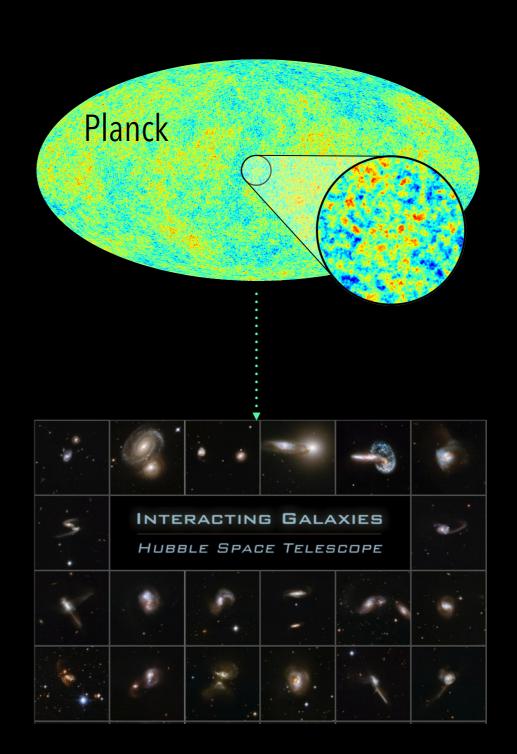




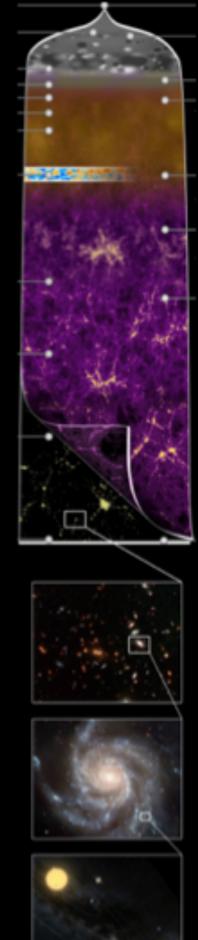


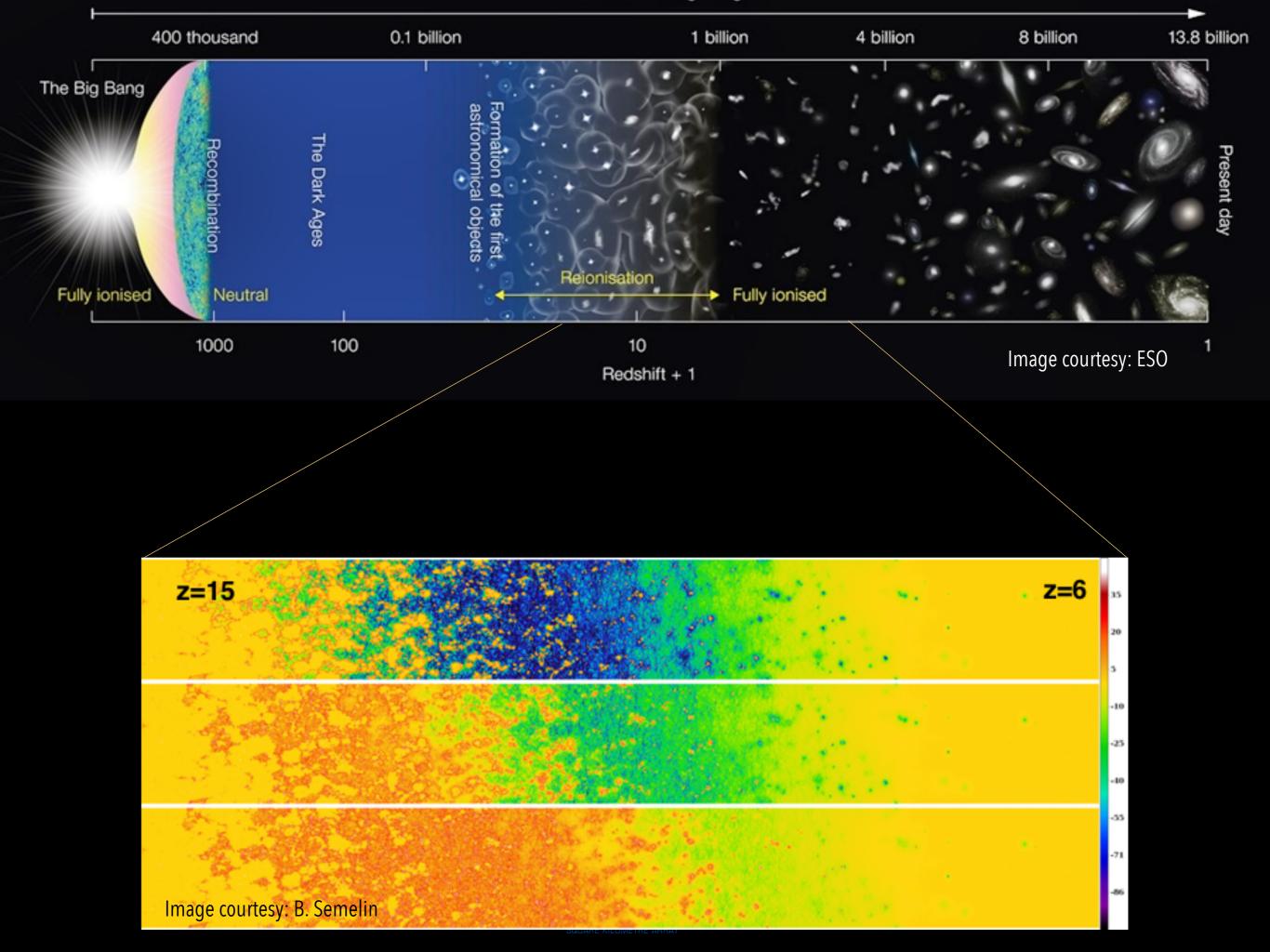


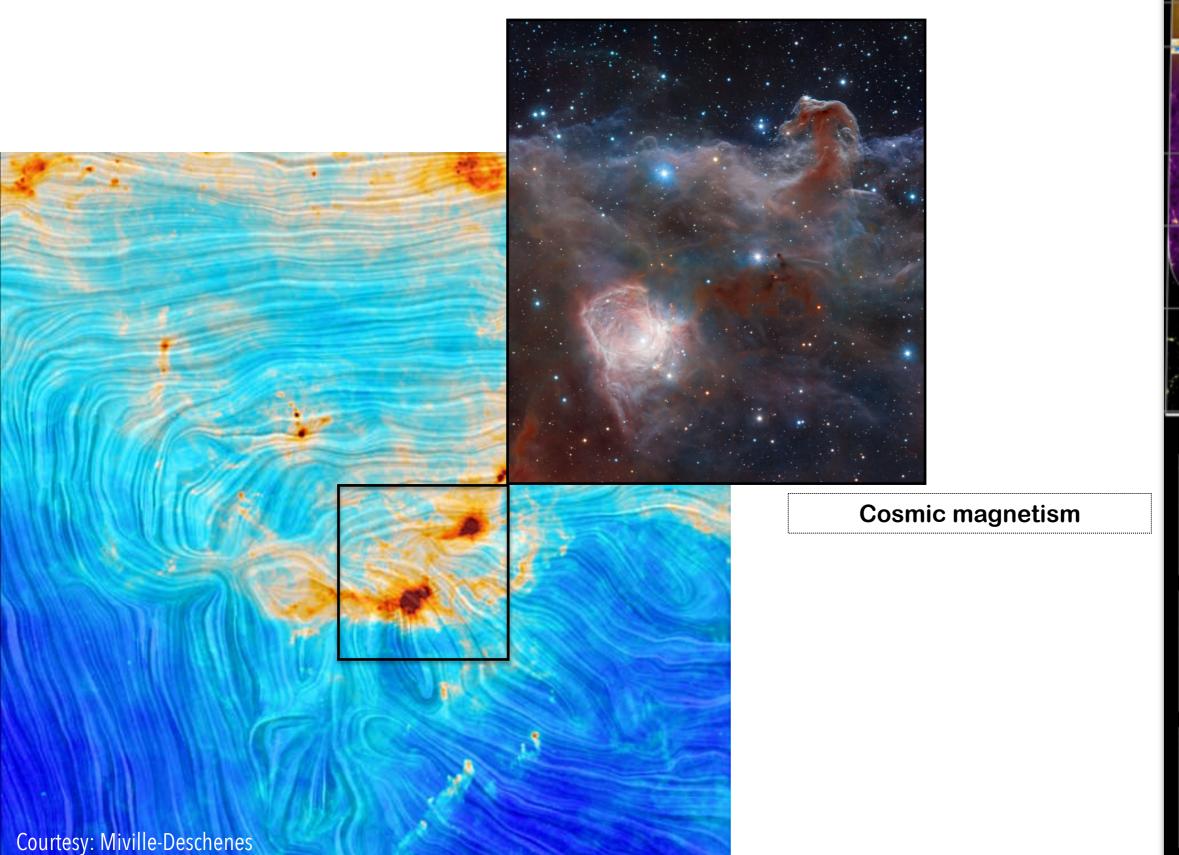


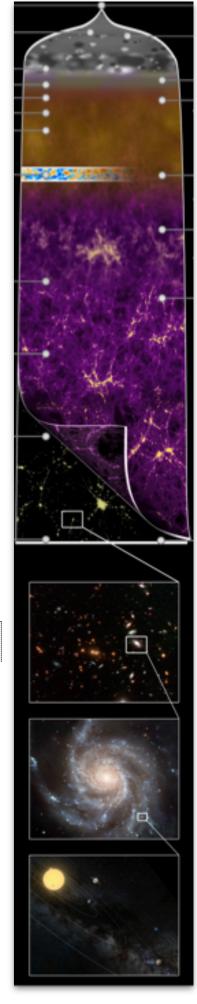






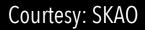




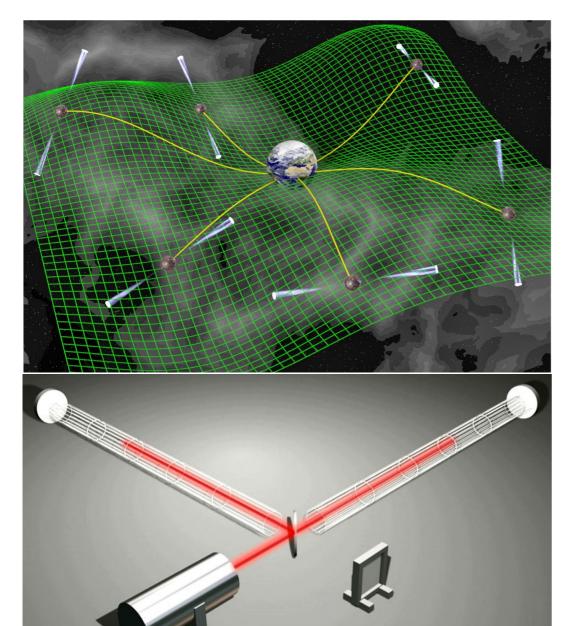


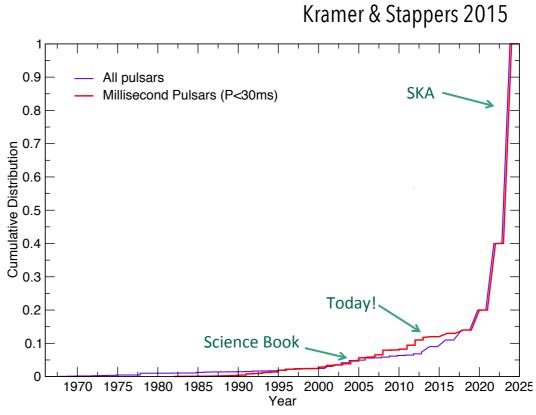


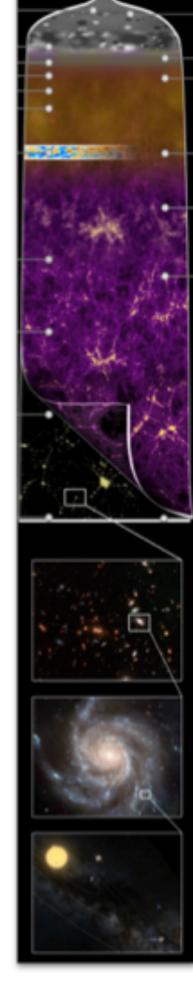






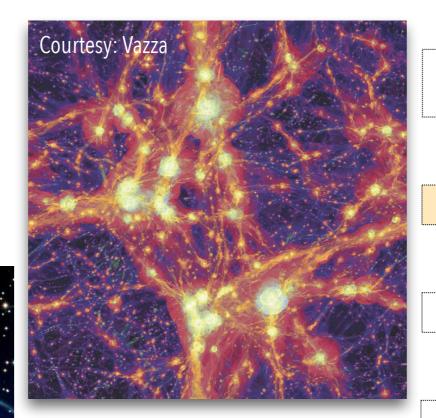








Courtesy: Calçada

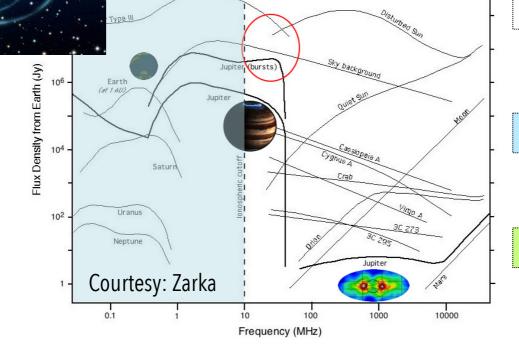


Cosmic dawn & Epoch of Reionisation

Cosmology

Galaxy evolution

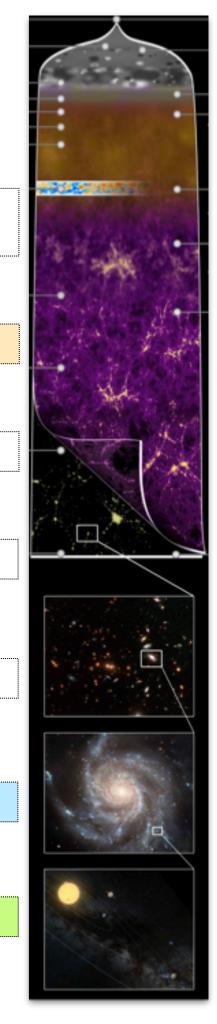
Cosmic magnetism



Fundamental physics

Transient sky

Cradle of life



Cosmology with the SKA

HI intensity mapping surveys:

• \sim 30,000 deg² out to $z\sim$ 3 in SKA1. Even though individual galaxies are not detected, the resolution is more than adequate to measure the fluctuations needed for BAO

Radio continuum surveys:

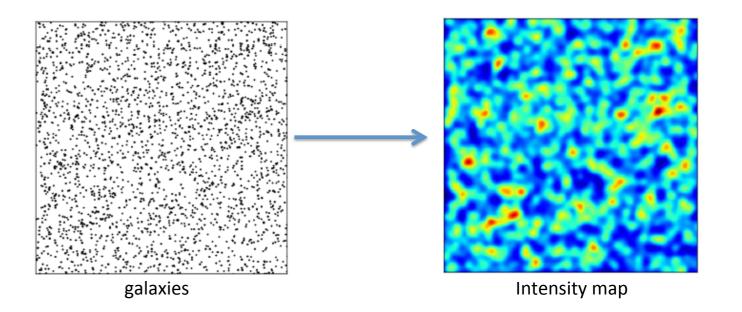
- \sim 30,000 deg² out to $z\sim$ 6, detecting \sim 10⁸ galaxies in SKA1
- ~10⁹ galaxies in SKA2

HI galaxy redshift surveys:

- \sim 5,000 deg² out to $z\sim$ 0.7, detecting \sim 10⁷ galaxies (SKA1)
- \sim 30,000 deg² out to $z\sim$ 2, detecting \sim 10⁹ galaxies (**SKA2 the billion galaxy survey**)

Weak lensing surveys:

- ~5,000deg² survey on SKA1
- ~30,000 deg² survey in SKA2



- Look at the total intensity for a given emission line in a large
 3d pixel (angle and frequency)
- Pixel will have joint emission from multiple galaxies

Cosmology with the SKA

HI intensity mapping surveys:

• \sim 30,000 deg² out to $z\sim$ 3 in SKA1. Even though individual galaxies are not detected, the resolution is more than adequate to measure the fluctuations needed for BAO

Radio continuum surveys:

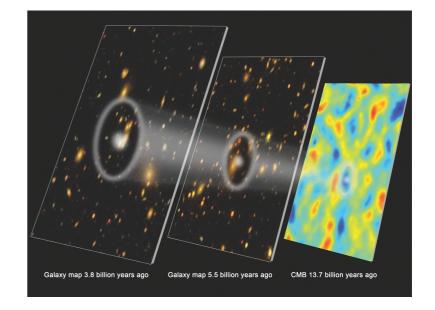
- \sim 30,000 deg² out to $z\sim$ 6, detecting \sim 10⁸ galaxies in SKA1
- ~10⁹ galaxies in SKA2

HI galaxy redshift surveys:

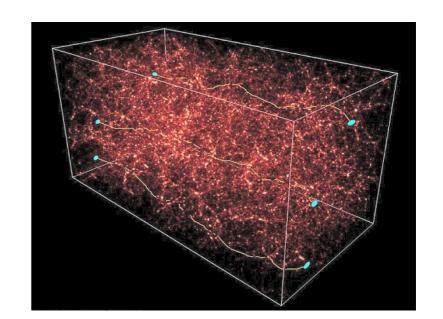
- \sim 5,000 deg² out to $z\sim$ 0.7, detecting \sim 10⁷ galaxies (SKA1)
- \sim 30,000 deg² out to $z\sim$ 2, detecting \sim 10⁹ galaxies (**SKA2 the billion galaxy survey**)

Weak lensing surveys:

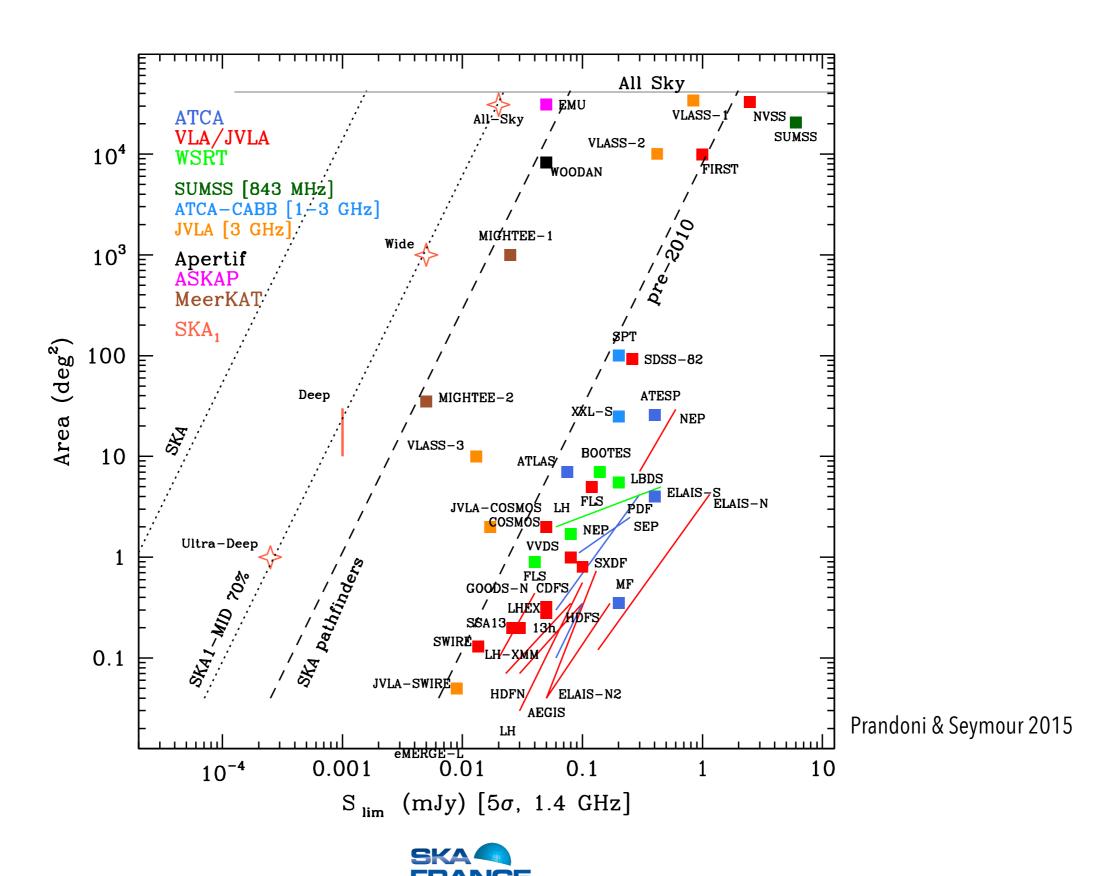
- ~5,000deg² survey on SKA1
- ~30,000 deg² survey in SKA2



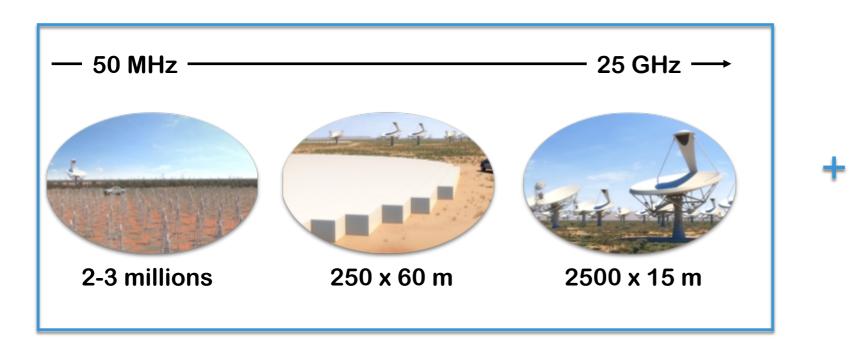
- the expansion history and geometry of the Universe
 - 'standard ruler' imprinted in the correlation function by the BAO
- ▶ the growth of structure
 - power spectrum of the observed density contrast
 - → weak lensing survey

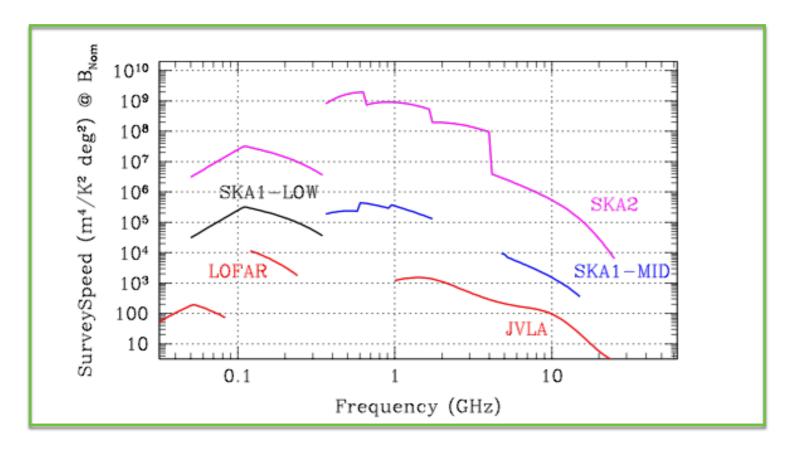


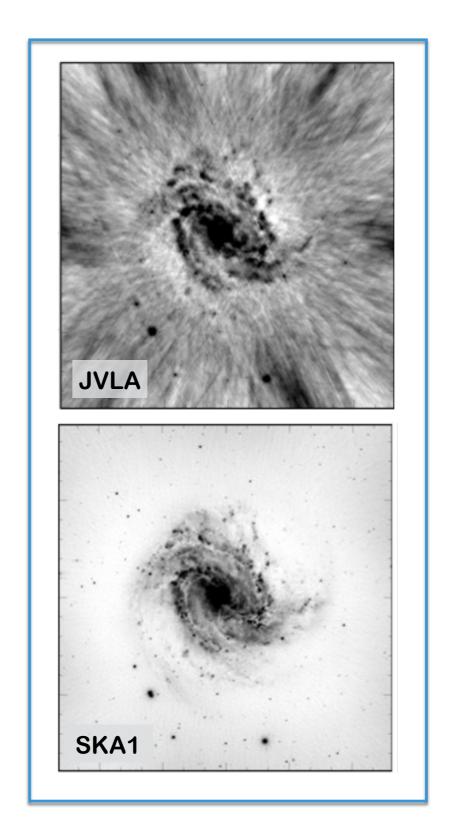
An exquisite instrument for surveying the sky



Why with the SKA









Observing modes

Non-imaging (Tide Array Beams)

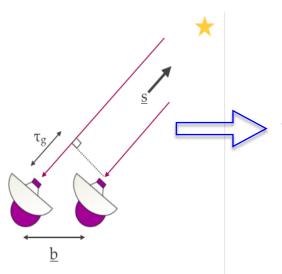
Pulsar Search & Timing

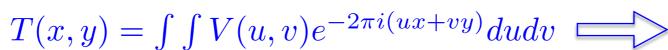
- Forms hundreds of beams within the dish/station beam
- Time resolution $\sim 60-100 \, \mu s$
- Data rate ~ 800 GB/s
- Archival ~ PB

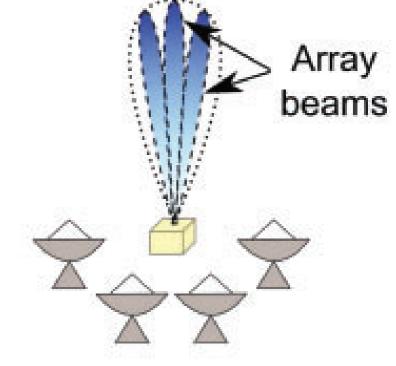
Imaging

Most of the other Key Science Topics

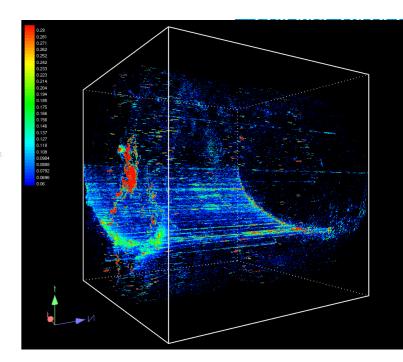
- 4D images (RA, Dec, Frequency, Polarization)
- Output image size ~ 1 PB
- Archival ~ 120 PB

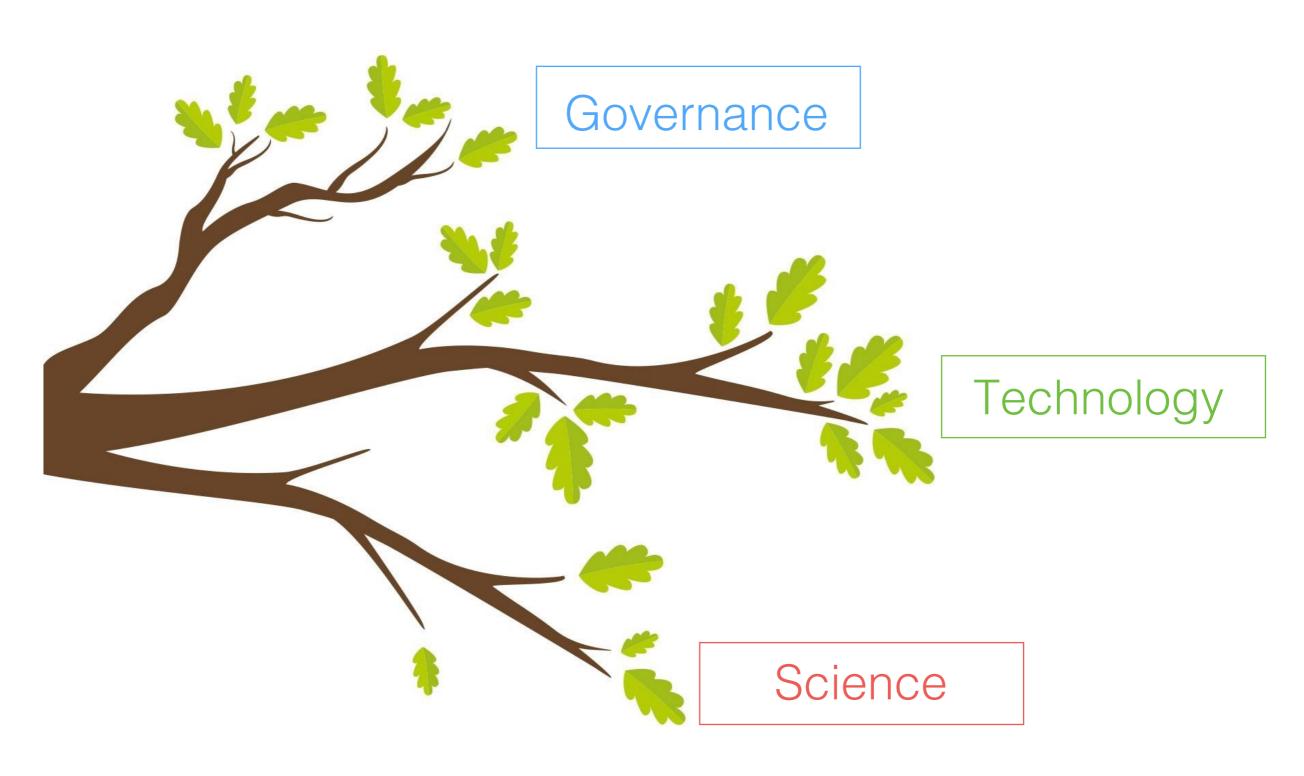












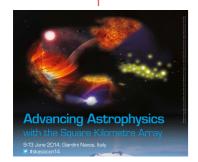




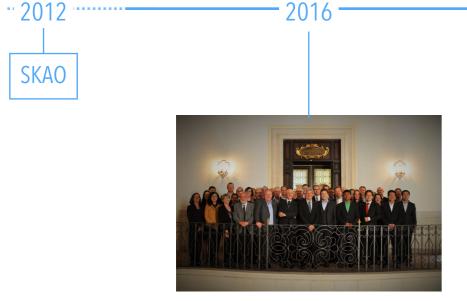












Technical consortia

Design Deployment Baseline Baseline SKA1-Mid 133 + 64* 130 + 64* No. dishes Max. Baseline 150 km 120 km Band 1 Feeds | 133 130 Band 2 Feeds 133 130 Band 5 Feeds 133 Pulsar Search 500 nodes 375 nodes (PSS) SKA1-Low 476 512 No. stations Max. Baseline 65 km 40 km Pulsar Search 167 nodes 125 nodes Common Compute Power | 260 PFLOPs 50 PFLOPs

SWG Meeting





SKA Phase 1 (SKA1)



SKA1-LOW (AUS) 130,000 log periodic antennas



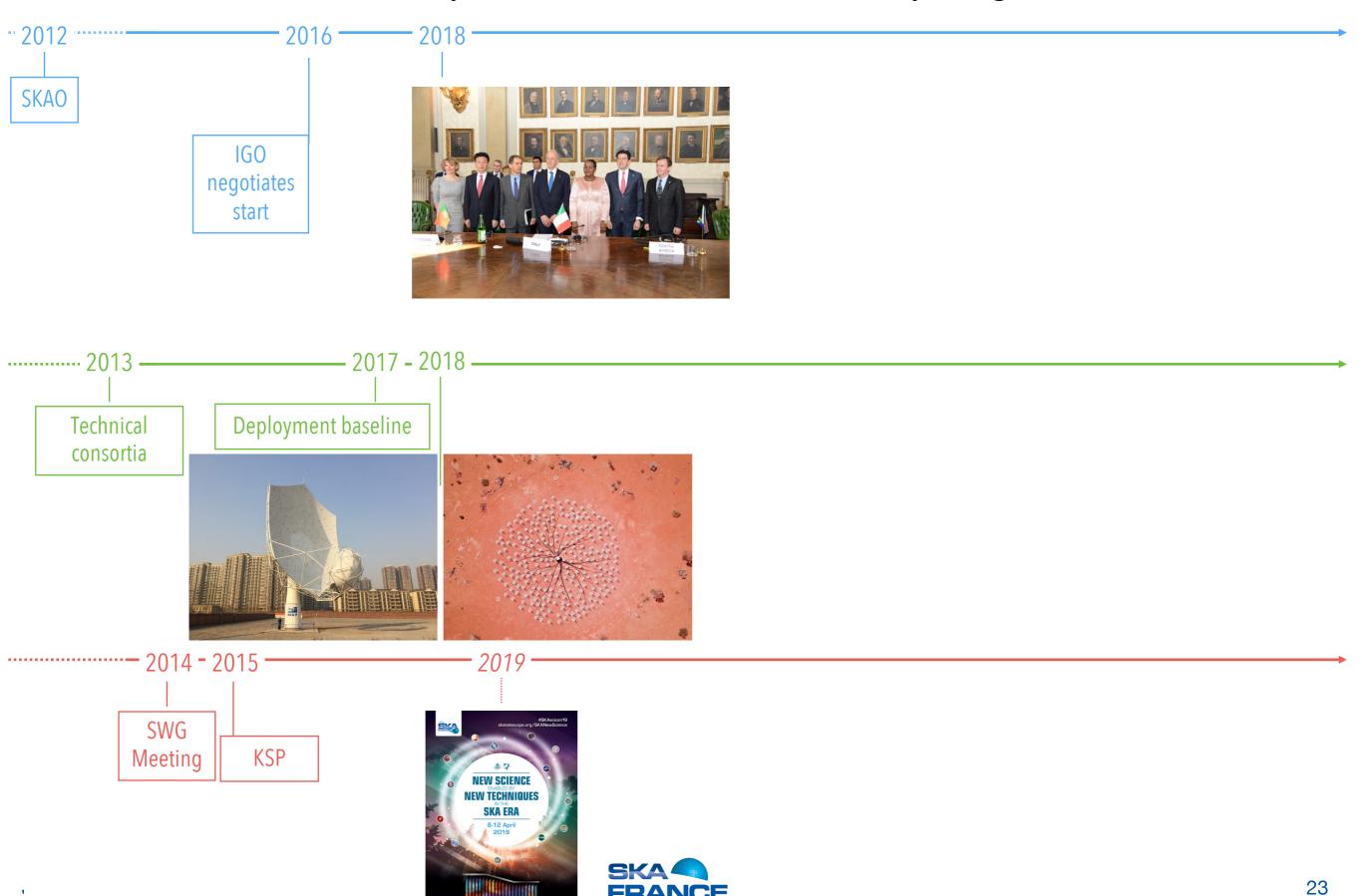
SKA1-MID (SA) 197 dishes (15m)

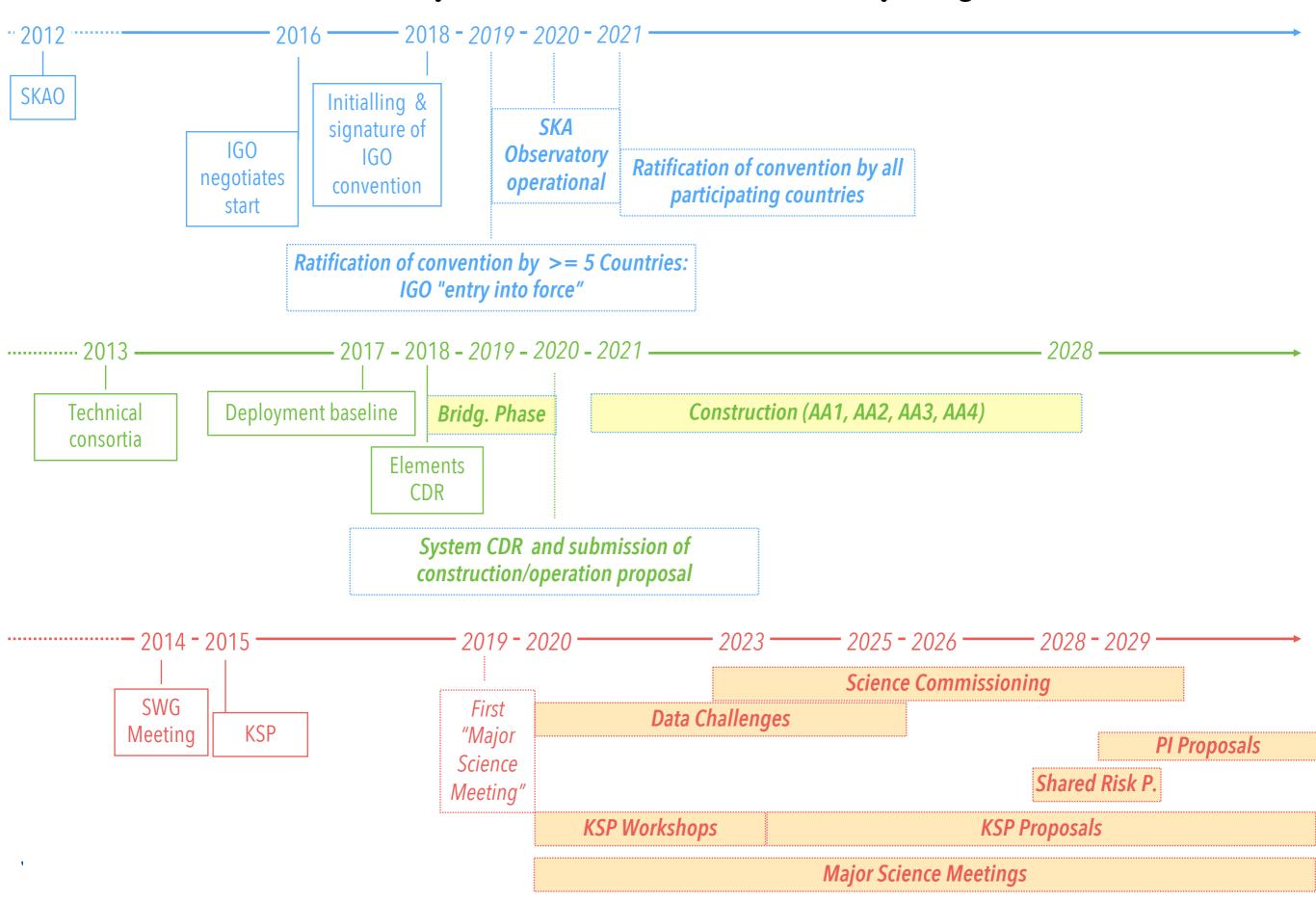
-50 MHz----------------------------------15 GHz→



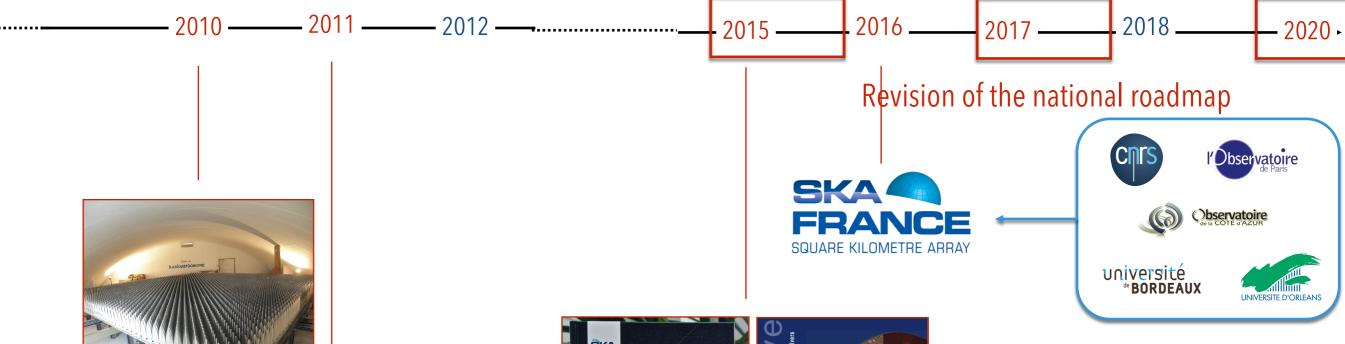








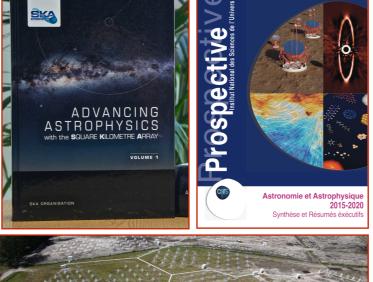
Timeline of the SKA project in France















The French SKA White Book

French SKA White Book

The French community towards the Square Kilometre Array



Editor in Chief:

C. Ferrari

Editors:

G. Lagache, J.-M. Martin, B. Semelin — Cosmology and Extra-galactic astronomy M. Alves, K. Ferrière, M.-A. Miville-Deschenes, L. Montier — Galactic Astronomy

- E. Josselin, N. Vilmer, P. Zarka Planets, Sun, Stars and Civilizations
- S. Corbel, S. Vergani Transient Universe
- S. Lambert, G. Theureau Fundamental Physics
- S. Bosse, A. Ferrari, S. Gauffre Technological Developments
- G. Marquette Industrial Perspectives and Solutions

178 authors from

* 40 French research institutes



* 6 private companies









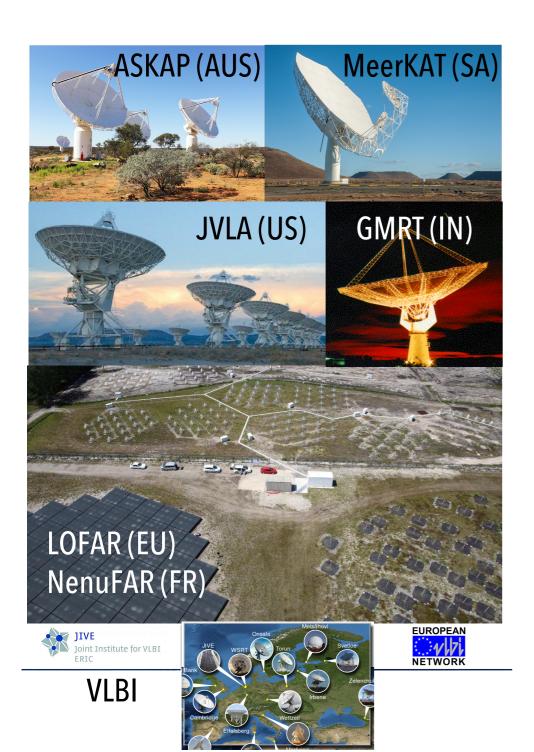






SKA: the radiotelescope of the XXIst century

SKA: a unique instrument



- French radio-astronomers are involved in the SKA scientific preparation
 - users of national et international radiotelescopes
 - all SKA Science Working Groups have French participants



A wider and wider community

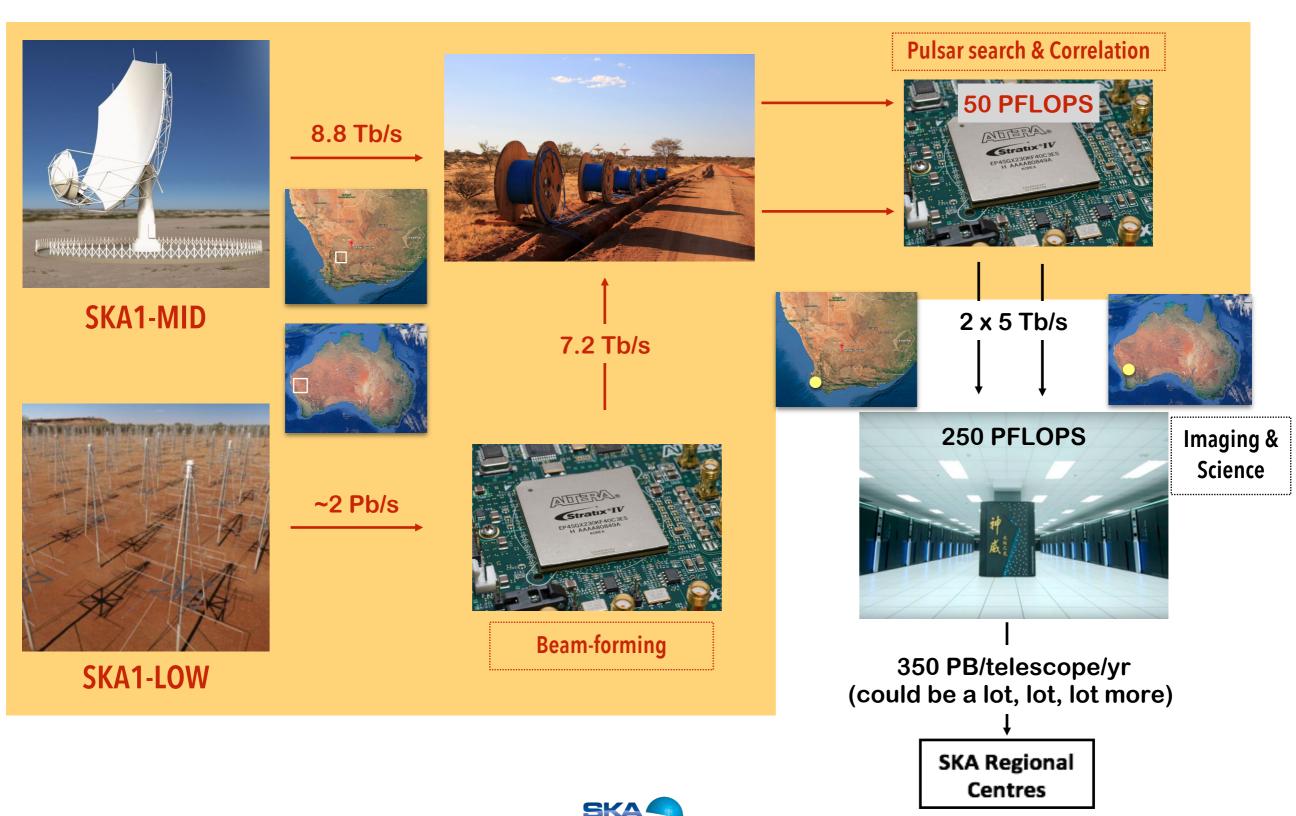
SKA: an interdisciplinary instrument



- The richest synergy chapter ever published about SKA vs. other projects, including:
 - instruments covering the whole electromagnetic spectrum
 - gravitational wave detectors



The SKA: an "exascale telescope"



February 2018: Maison SKA France





Bull







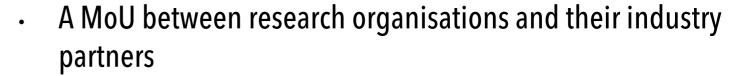


Air Liquide









- A science and technology roadmap
- A forum to develop fundamental research and R&D projects







July 2018: CNRS new member of SKA Organisation

February 2018: Maison SKA France



université BORDEAUX















- A MoU between research organisations and their industry partners
- A science and technology roadmap
- A forum to develop fundamental research and R&D projects













July 2018: CNRS new member of SKA Organisation



Thanks!

chiara.ferrari@oca.eu