



Galileo, 1610

PHYSICS with CURRENT and FUTURE GW EXPERIMENTS

Giovanni Losurdo –

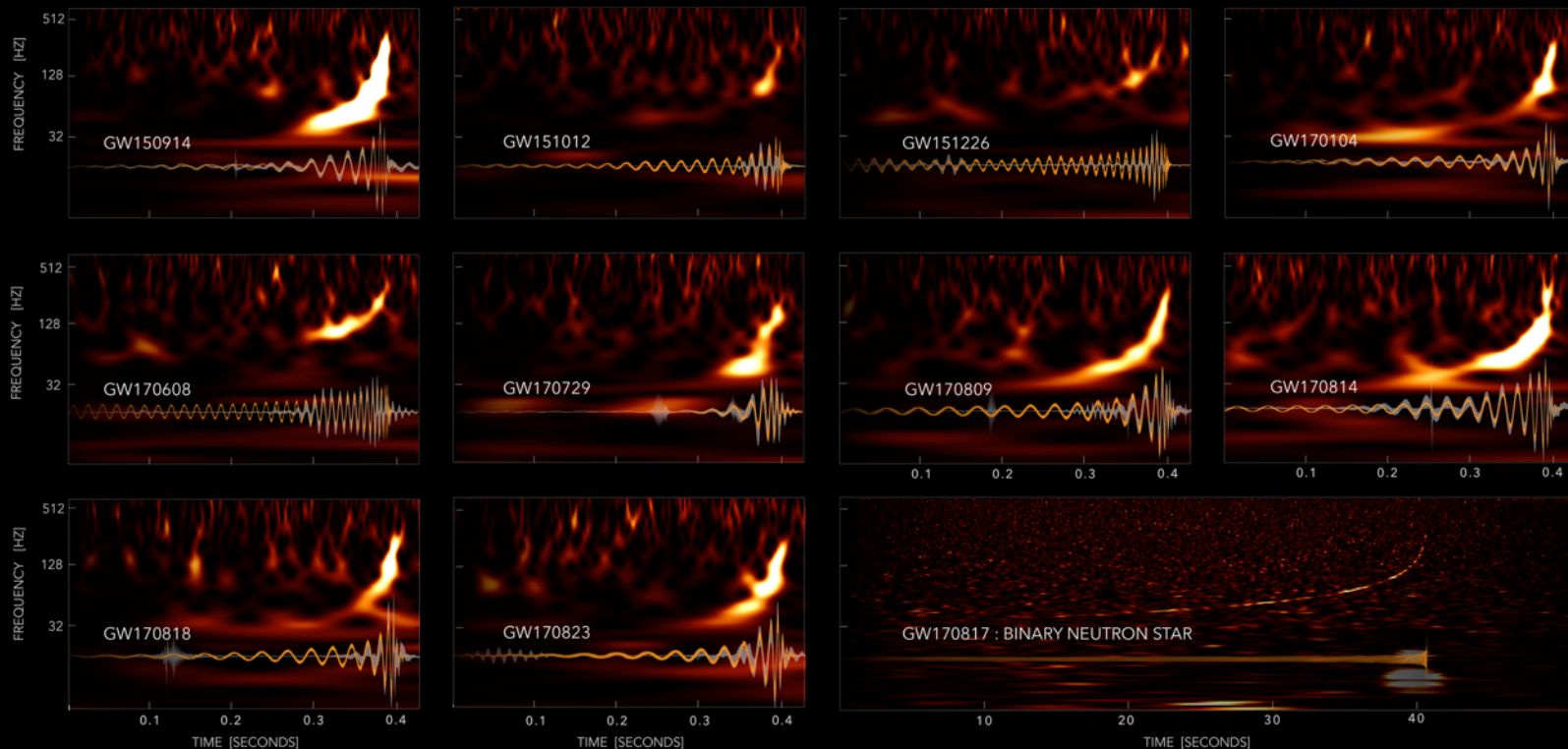


Pisa

Slide credits: E Chassand-Mottin, P Schmidt, G Gemme, M Punturo



GRAVITATIONAL-WAVE TRANSIENT CATALOG-1



LIGO-VIRGO DATA: [HTTPS://DOI.ORG/10.7935/B2H3-HH23](https://doi.org/10.7935/b2h3-hh23)

WAVELET (UNMODELED) EINSTEIN'S THEORY

S. GHONGE, K. JANI | GEORGIA TECH

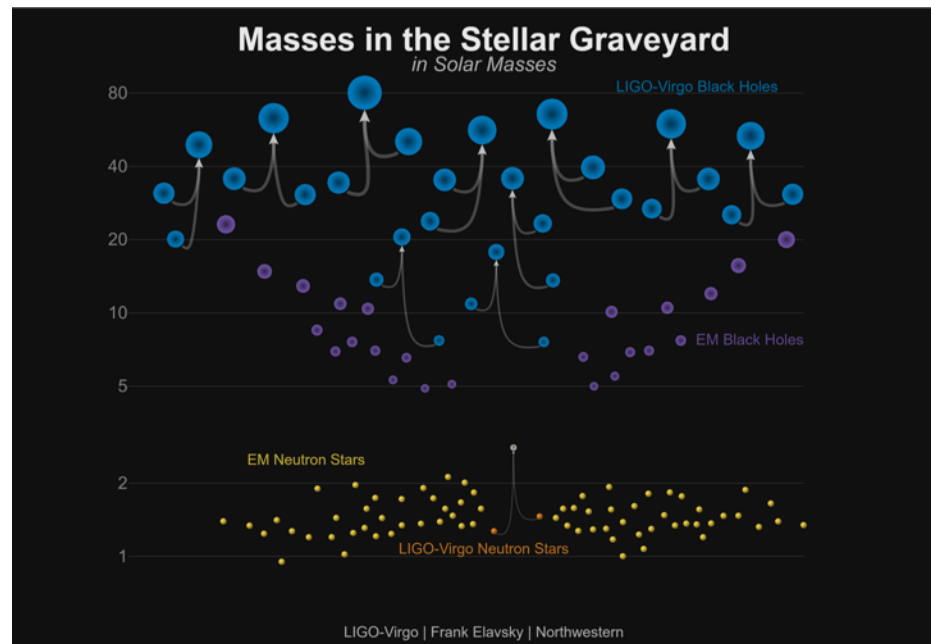
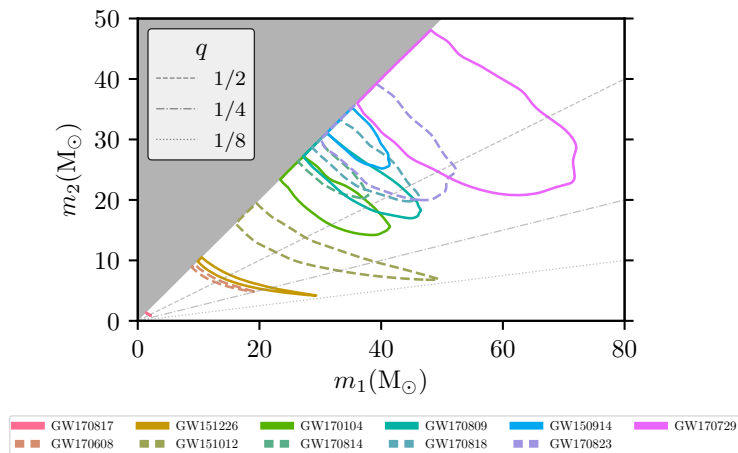
The Catalog

B. P. Abbott, et al. (LIGO Virgo Collaboration), PRX,
9, 031040 (2019)

Information on masses, spins, energy radiated, position, distance, inclination, polarization.
Population distribution may shed light on formation mechanisms

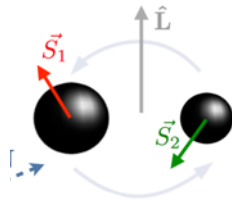
Event	$m_1/M_\odot$	$m_2/M_\odot$	$M/M_\odot$	χ_{eff}	$M_f/M_\odot$	a_f	$E_{\text{rad}}/(M_\odot c^2)$	$\ell_{\text{peak}}/(\text{erg s}^{-1})$	D_L/Mpc	z	$\Delta\Omega/\text{deg}^2$
GW150914	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	$-0.01^{+0.12}_{-0.13}$	$63.1^{+3.3}_{-3.0}$	$0.69^{+0.05}_{-0.04}$	$3.1^{+0.4}_{-0.4}$	$3.6^{+0.4}_{-0.4} \times 10^{56}$	430^{+150}_{-170}	$0.09^{+0.03}_{-0.03}$	194
GW151012	$23.2^{+14.0}_{-5.4}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.2}$	$0.04^{+0.28}_{-0.19}$	$35.7^{+9.9}_{-3.7}$	$0.67^{+0.13}_{-0.11}$	$1.5^{+0.5}_{-0.5}$	$3.2^{+0.8}_{-1.7} \times 10^{56}$	1060^{+540}_{-480}	$0.21^{+0.09}_{-0.09}$	1491
GW151226	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	$0.18^{+0.20}_{-0.12}$	$20.5^{+6.4}_{-1.5}$	$0.74^{+0.07}_{-0.05}$	$1.0^{+0.1}_{-0.2}$	$3.4^{+0.7}_{-1.7} \times 10^{56}$	440^{+180}_{-190}	$0.09^{+0.04}_{-0.04}$	1075
GW170104	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	$-0.04^{+0.17}_{-0.20}$	$49.4^{+5.2}_{-3.9}$	$0.66^{+0.09}_{-0.11}$	$2.2^{+0.5}_{-0.5}$	$3.2^{+0.7}_{-1.0} \times 10^{56}$	960^{+430}_{-410}	$0.19^{+0.07}_{-0.08}$	912
GW170608	$11.2^{+5.4}_{-1.9}$	$7.5^{+1.5}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	$0.04^{+0.19}_{-0.06}$	$17.9^{+3.4}_{-0.7}$	$0.69^{+0.04}_{-0.04}$	$0.8^{+0.1}_{-0.1}$	$3.4^{+0.5}_{-1.3} \times 10^{56}$	320^{+120}_{-110}	$0.07^{+0.02}_{-0.02}$	524
GW170729	$50.7^{+16.3}_{-10.2}$	$34.4^{+8.9}_{-10.2}$	$35.8^{+6.3}_{-4.9}$	$0.37^{+0.21}_{-0.26}$	$80.3^{+14.5}_{-10.3}$	$0.81^{+0.07}_{-0.13}$	$4.9^{+1.6}_{-1.7}$	$4.2^{+0.8}_{-1.5} \times 10^{56}$	2760^{+1290}_{-1350}	$0.48^{+0.18}_{-0.21}$	1069
GW170809	$35.2^{+8.3}_{-5.9}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	$0.07^{+0.17}_{-0.16}$	$56.4^{+5.2}_{-3.7}$	$0.70^{+0.08}_{-0.09}$	$2.7^{+0.6}_{-0.6}$	$3.5^{+0.6}_{-0.9} \times 10^{56}$	990^{+320}_{-380}	$0.20^{+0.05}_{-0.07}$	310
GW170814	$30.7^{+5.5}_{-2.9}$	$25.6^{+2.8}_{-4.0}$	$24.3^{+1.4}_{-1.1}$	$0.07^{+0.12}_{-0.11}$	$53.6^{+3.2}_{-2.5}$	$0.73^{+0.07}_{-0.05}$	$2.8^{+0.4}_{-0.3}$	$3.7^{+0.5}_{-0.5} \times 10^{56}$	560^{+140}_{-210}	$0.12^{+0.03}_{-0.04}$	99
GW170817	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	$0.00^{+0.02}_{-0.01}$	≤ 2.8	≤ 0.89	≥ 0.04	$\geq 0.1 \times 10^{56}$	40^{+10}_{-10}	$0.01^{+0.00}_{-0.00}$	22
GW170818	$35.5^{+7.5}_{-4.7}$	$26.9^{+4.4}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	$-0.09^{+0.18}_{-0.21}$	$59.8^{+4.8}_{-3.7}$	$0.67^{+0.07}_{-0.08}$	$2.7^{+0.5}_{-0.5}$	$3.4^{+0.5}_{-0.7} \times 10^{56}$	1020^{+430}_{-370}	$0.20^{+0.07}_{-0.07}$	35
GW170823	$39.5^{+10.1}_{-6.6}$	$29.4^{+6.5}_{-7.1}$	$29.3^{+4.2}_{-3.1}$	$0.08^{+0.19}_{-0.22}$	$65.6^{+9.3}_{-6.5}$	$0.71^{+0.08}_{-0.09}$	$3.3^{+0.9}_{-0.8}$	$3.6^{+0.6}_{-0.9} \times 10^{56}$	1860^{+840}_{-840}	$0.34^{+0.13}_{-0.14}$	1780

Masses

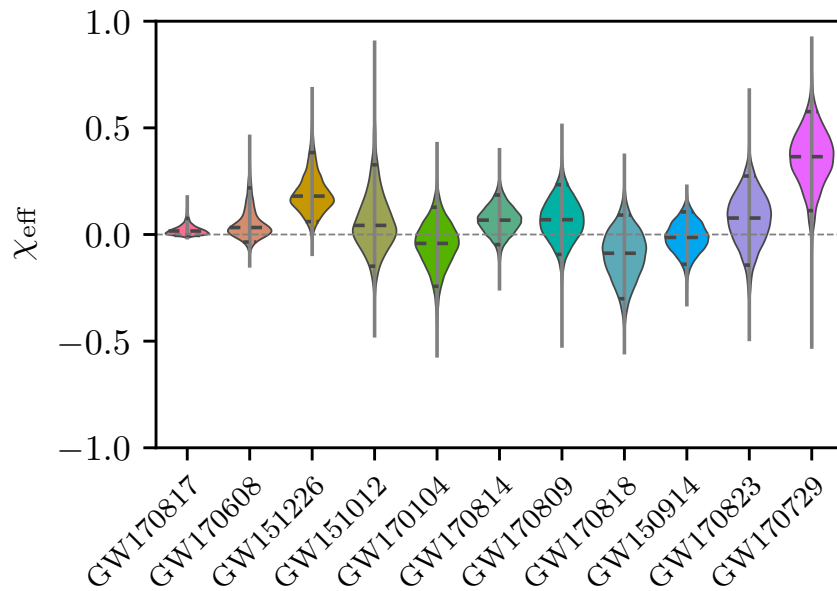


- ❑ Wide range of BH masses
 - More BBH towards the heavy end, none in either mass gaps
 - GW170729: heaviest ($\sim 80 M_\odot$) and furthest ($z \sim 0.5$)
- ❑ BNS remnant in the mass gap

BH Spins



$$\chi_{\text{eff}} = \frac{(m_1 \vec{\chi}_1 + m_2 \vec{\chi}_2) \cdot \hat{L}_N}{M}$$

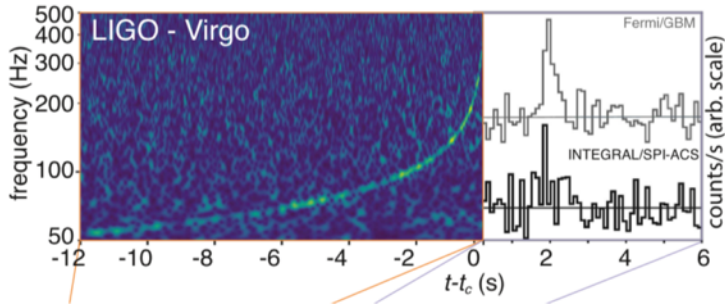


- Key signatures to discriminate BH populations: shed light on formation mechanism
- GW170729 has a clear indication of a net positive χ_{eff}
- Precession remains unconstrained for all events

GW170817

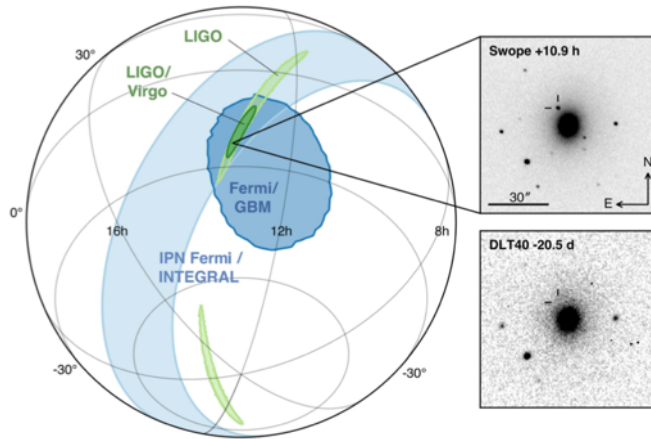
ApJL 2017 and ref. therein

Multiwavelength observation
of the source: kilonova



Aug 17, 2017

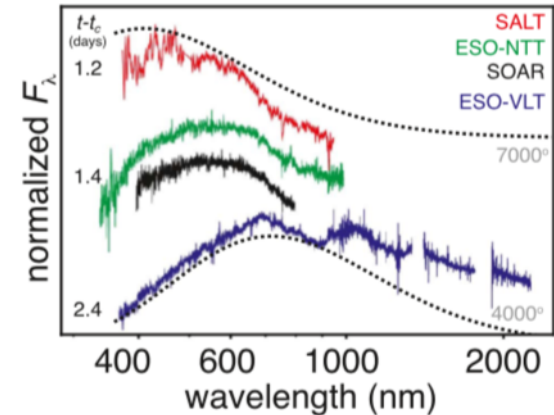
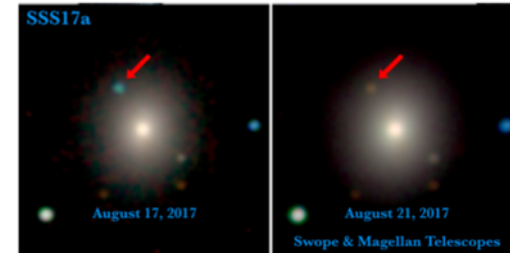
GW from BNS detected by LIGO/Virgo
Fermi GBM detects sGRB after 1.7 s



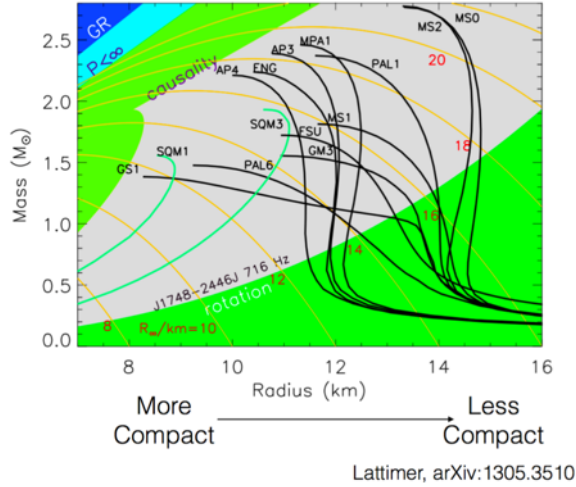
GCN sent after $\sim 27'$

Event localized within $\sim 30^\circ$ sq

Optical counterpart first observed by
SWOPE after ~ 11 h

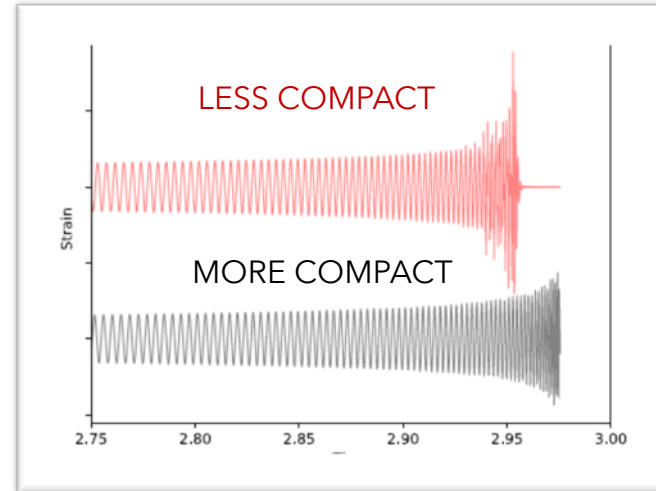
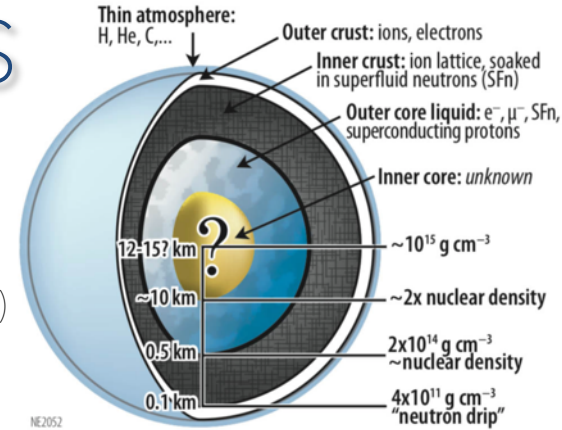


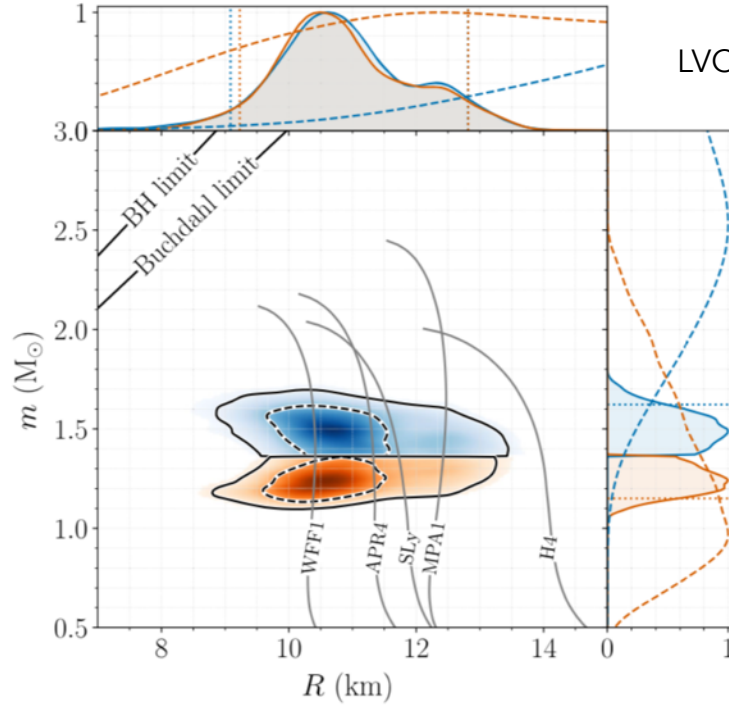
GW170817: NUCLEAR PHYSICS



INTERNAL STRUCTURE AND COMPOSITION (LARGELY UNKNOWN) ENCODED IN THE EQUATION OF STATE

Each NS is deformed by the tidal field of the companion
Deformations leave an imprint on the GW emission

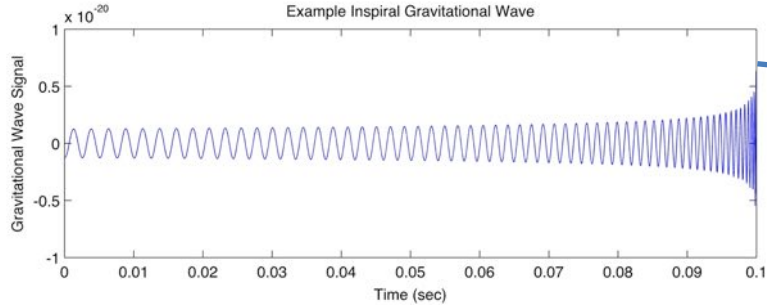




LVC, PRL 121, 161101 (2018)

LIGO/VIRGO are measuring the NS radius/mass. Improving the accuracy (better SNR, more events) will allow to determine the EOS

GW170817: COSMOLOGY



$$H_0 d_L = cz$$

$$H_0 = 70_{-8.0}^{+12.0} \text{ km s}^{-1} \text{ Mpc}$$



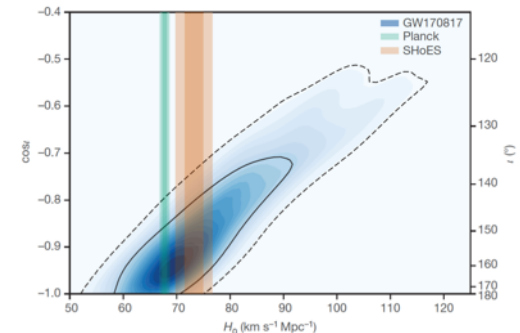
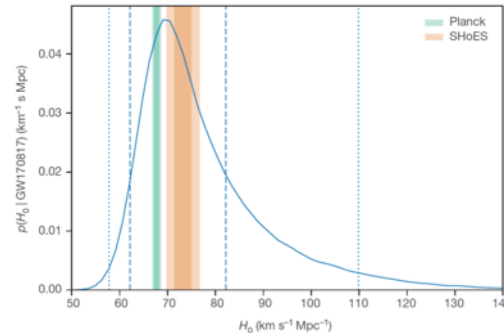
From GW: redshifted mass and luminosity distance

$$\Psi(f) \propto 2\pi f t_c - \Phi_c - \frac{\pi}{4} + \frac{3}{128} (\pi M_z f)^{-5/3} [1 + \dots]$$

$$\tilde{h}_+(f) \propto A_0 \frac{M_z^{5/6}}{D_L} [1 + \cos^2 \iota] f^{-7/6} e^{i\Psi(f)}$$

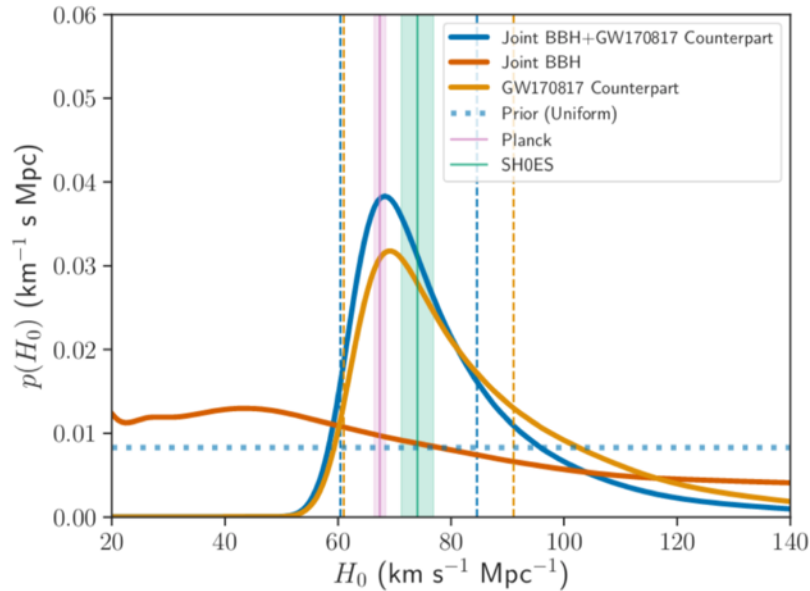
Inclination angle introduces degeneracy, which will be removed by measuring the 2 polarizations

Abbott et al., Nature, 24471 (2017)



New measurement of H_0 using the O1+O2 detections and galaxy catalogs

[arxiv:1908.06060](https://arxiv.org/abs/1908.06060)



$$H_0 = 68_{-7}^{+14} \text{ km s}^{-1} \text{Mpc}^{-1}$$

GW170817: FUNDAMENTAL PHYSICS

- GW170817 provides a stringent test of the speed of gravitational waves

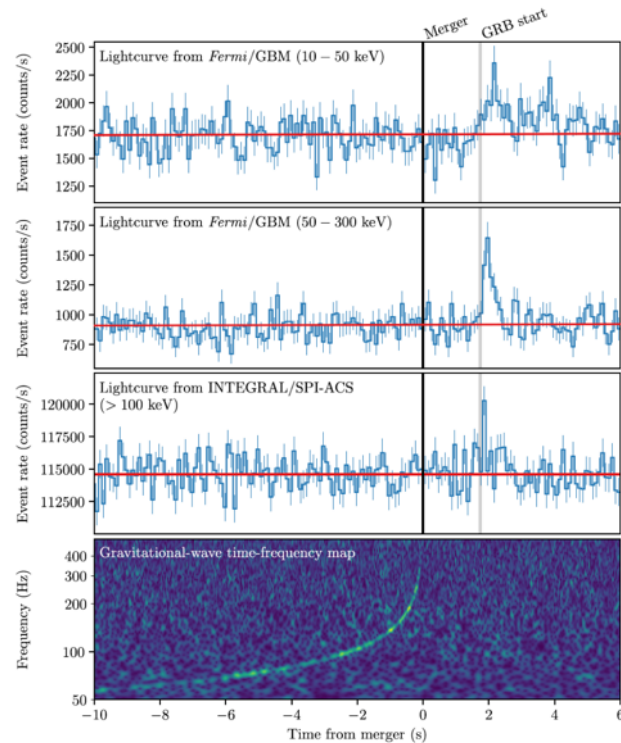
$$\frac{v_{GW} - c}{c} \approx \frac{c}{D} (\Delta t_{obs} - \Delta t_{start})$$

- $\Delta t = 1.7 \pm 0.5$ s over ~ 130 Myrs

$$-3 \cdot 10^{-15} \leq \frac{v_{GW} - c}{c} \leq 7 \cdot 10^{-16}$$

Assuming $D=26$ Mpc (lowest value)

Assuming sGRB emitted
10 s after GW



Abbott et al., ApJL., 848:L13 (2017)

More science from O1-O2

- ❑ Updated **merger rates**: LVC, GWTC-1, PRX 2019
 - BNS: $[110, 3840] \text{ Gpc}^{-3} \text{ y}^{-1}$
 - BBH: $[9.7, 101] \text{ Gpc}^{-3} \text{ y}^{-1}$
 - NSBH merger rate 90% upper limit of $610 \text{ Gpc}^{-3} \text{ y}^{-1}$
- ❑ **Propagation of GW** (dispersion): $m_g \leq 4.7 \times 10^{-23} \text{ eV}/c^2$ LVC, arXiv 1903.04467
- ❑ **Polarization** tests LVC, PRL 2017
- ❑ Tests of the nature of the BNS **remnant** LVC, arXiv 1908.01012
- ❑ Data available: <https://www.gw-openscience.org/catalog>

Getting Started

Data

Catalogs

Bulk Data

Tutorials

Software

Detector Status

Timelines

My Sources

GPS ↔ UTC

About the detectors

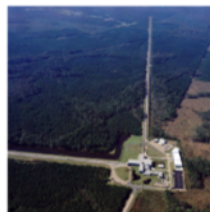
Projects

Acknowledge

GWOSC



LIGO Hanford Observatory, Washington
(Credits: C. Gray)



LIGO Livingston Observatory, Louisiana
(Credits: J. Giaime)



Virgo detector, Italy
(Credits: Virgo Collaboration)

The Gravitational Wave Open Science Center provides data from **gravitational-wave observatories, along with access to **tutorials** and **software tools**.**



Get started!



Download data



GWTC-1: Catalog of Compact Binary Mergers



Join the email list



Explore the open data web course

LIGO and Virgo's portal for

- Bulk data
- Event 1-hour time-series data, etc.
- Pointers to papers, data behind figures, posterior samples
- Pointers to analysis codes
- Pointers to Workshop materials



The O3 run

O3a: Apr 1 2019 → Oct 1 2019

O3b: Nov 1 2019 → May 1 2020 (with KAGRA)

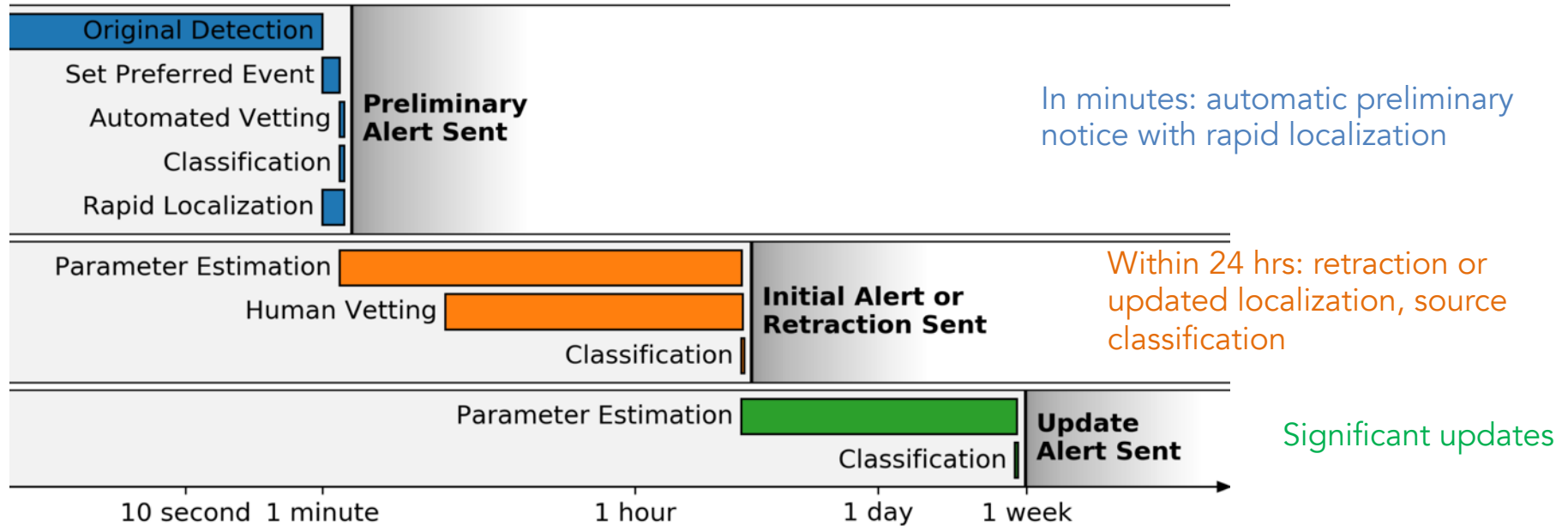


Open Public Alerts



LIGO-Virgo will issue Open Public Alerts during the O3 run

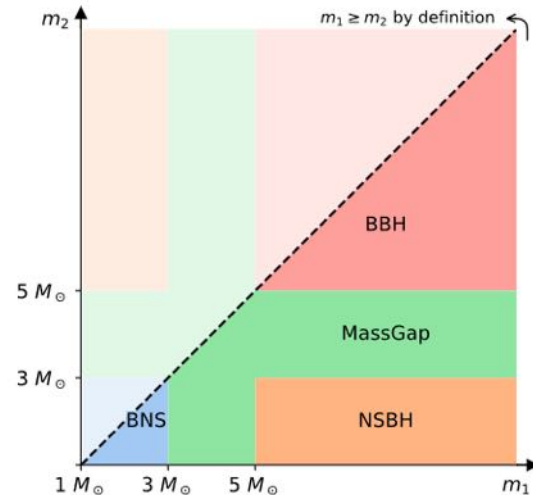
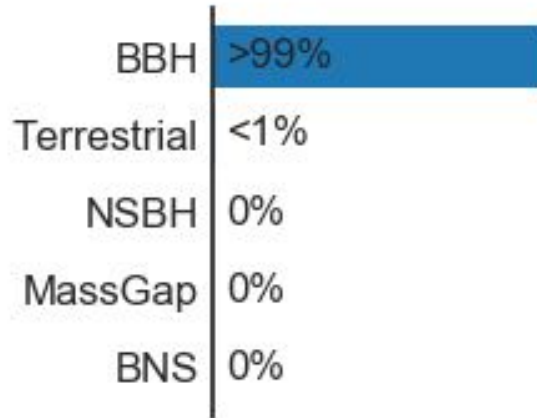
Time since gravitational-wave signal



Open Public Alerts

- ❑ Localization: 3D map for follow-up
- ❑ Classification: Five numbers, summing to unity, giving probability that the source belongs to five categories
 - This assumes that terrestrial and astrophysical events occur as independent Poisson processes

[arXiv:1903.06881](https://arxiv.org/abs/1903.06881)



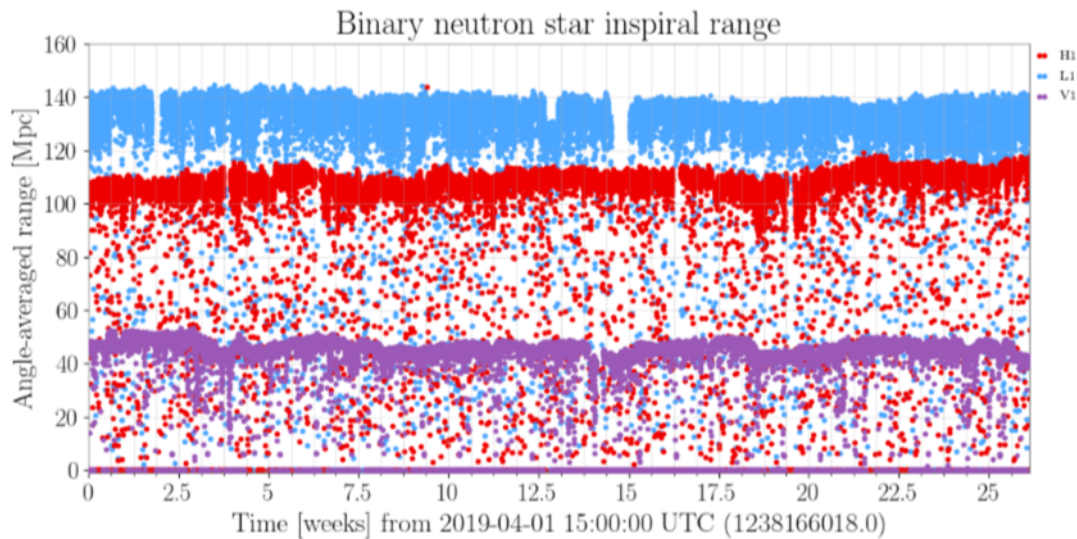
O3



Network duty factor

[1238166018-1259193618]

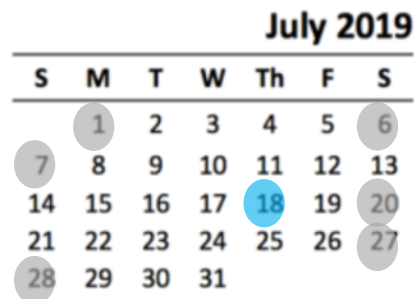
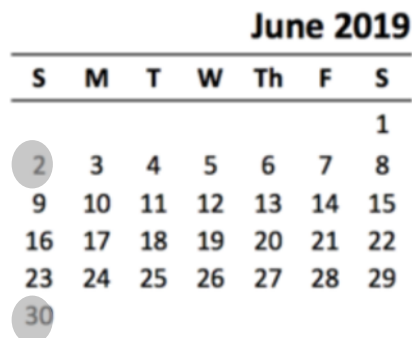
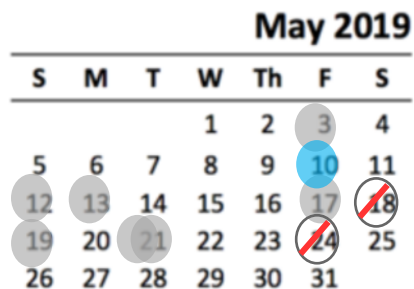
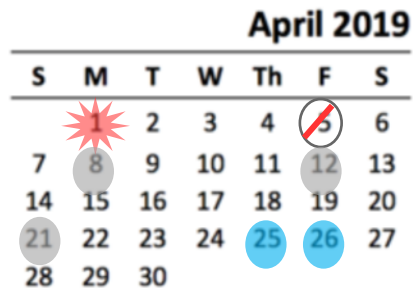
- Triple interferometer [44.1%]
- Double interferometer [37.7%]
- Single interferometer [15.1%]
- No interferometer [3.2%]



- BBH
- BNS
- NSBH
- Mass Gap

O3a - Summary of public alerts

 Retracted

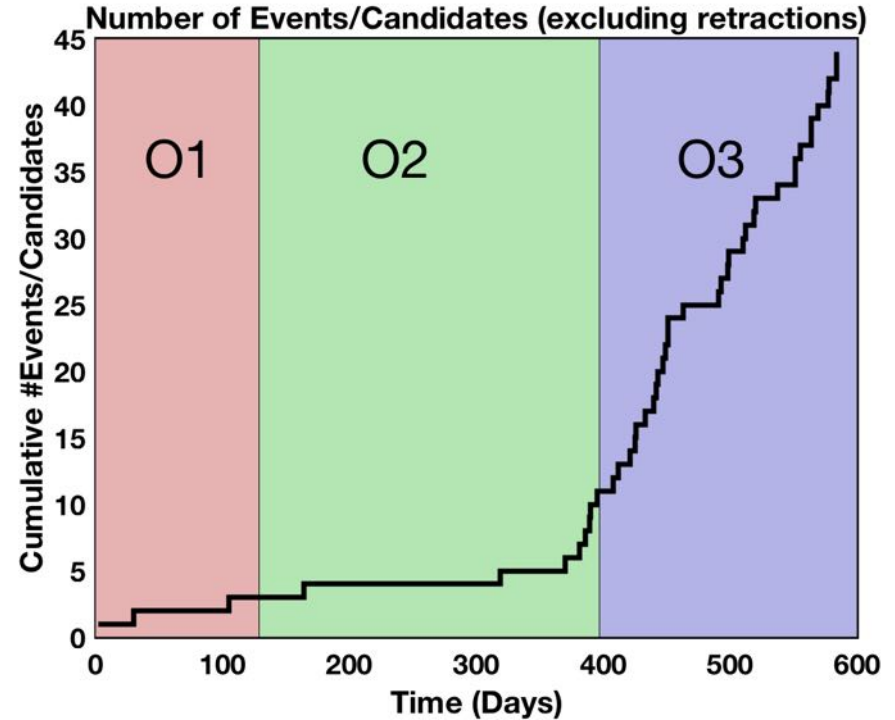


Observation Run	Network	Expected BNS Detections	Expected NSBH Detections	Expected BBH Detections
O3	HLV	2^{+8}_{-2}	0^{+19}_{-0}	15^{+19}_{-10}
O4	HLVK	8^{+42}_{-7}	2^{+94}_{-2}	68^{+81}_{-38}

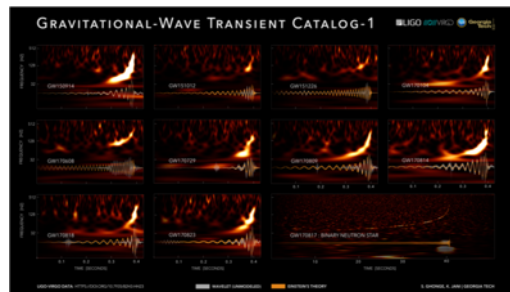
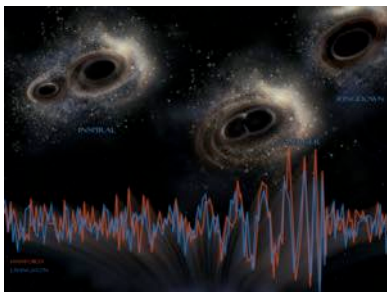
<https://arxiv.org/abs/1304.0670>

- ❑ 33 public alerts (excluding retractions) in O3a
- ❑ Most interesting:
 - 3 NSBH ($p > 50\%$)
 - 4 BNS ($p \gtrsim 50\%$)
 - 1 system with a star in the "mass gap"
 - 1 BBH at $z \gtrsim 1$
- ❑ Extensive follow-up but no e.m. counterpart in O3a

<https://gracedb.ligo.org/latest/>

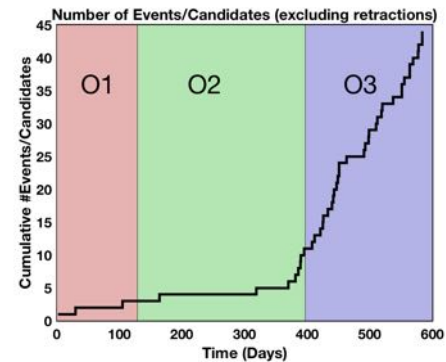
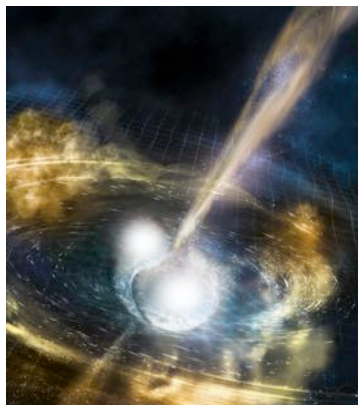


In just 4 years...



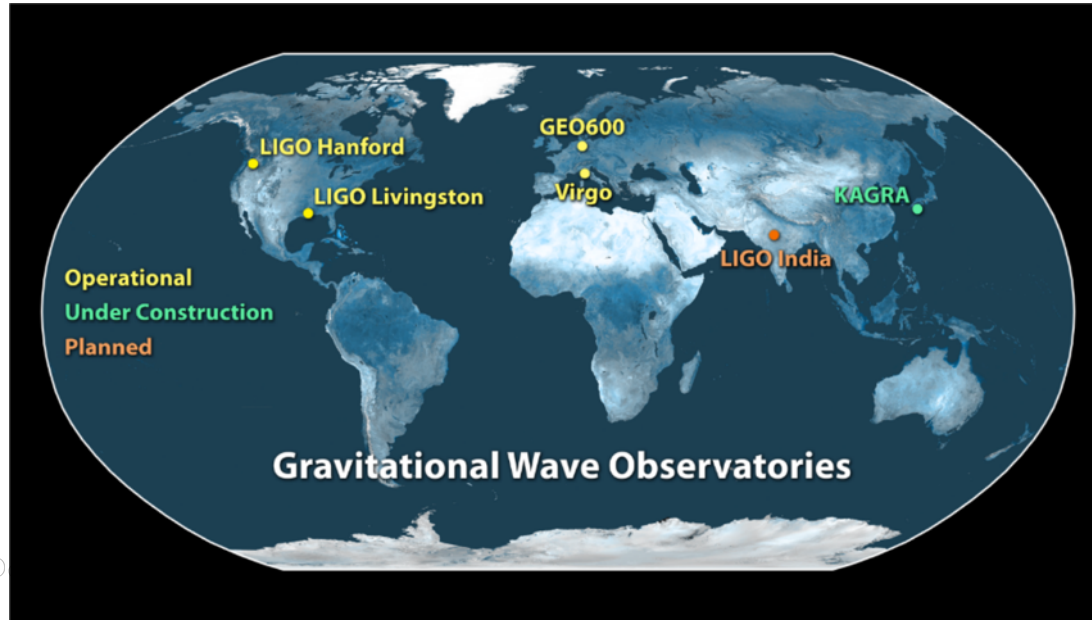
...to population studies

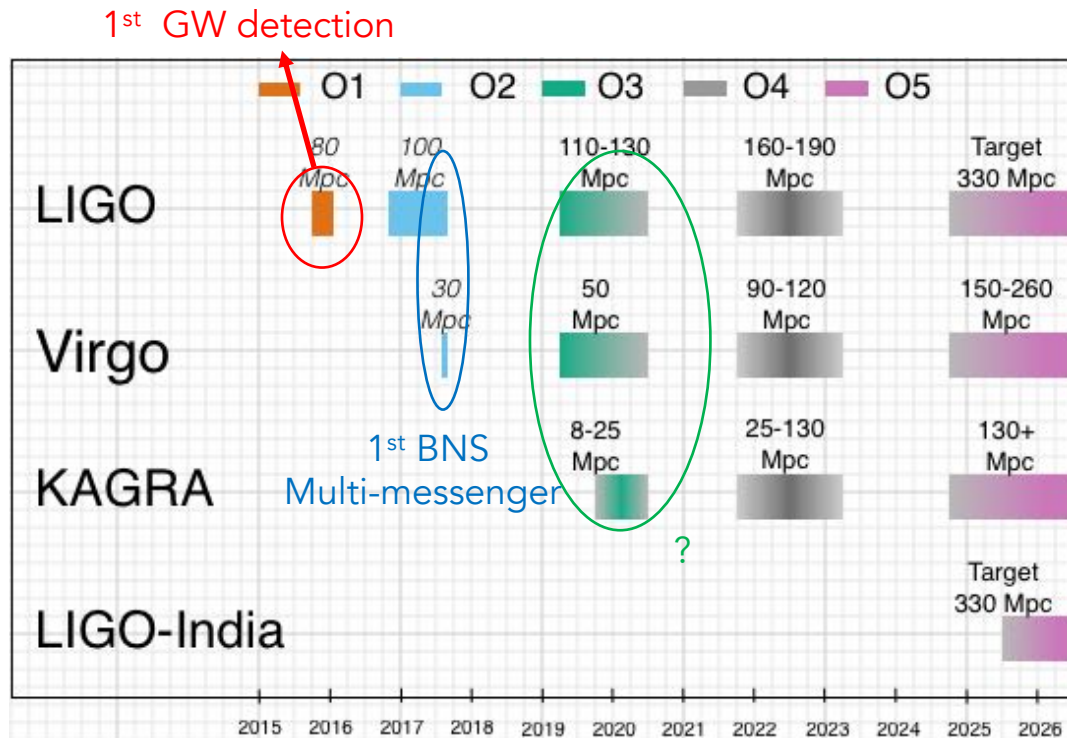
From discovery...



MID-TERM OUTLOOK

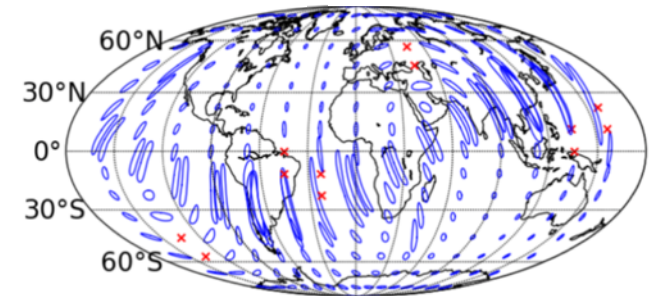
- ❑ KAGRA to join the network in O3b
- ❑ Ligo/Virgo to be further upgraded (A+/AdV+) after O3
- ❑ Ligo India to join the network in ~2024



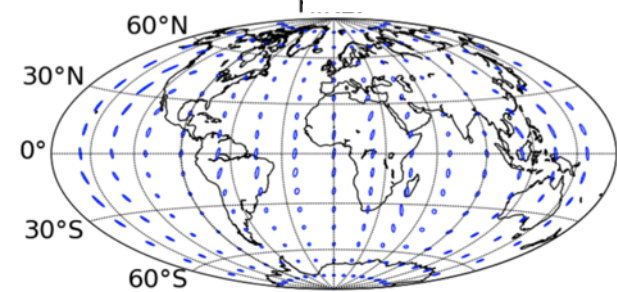


[arXiv:1304.0670](https://arxiv.org/abs/1304.0670) (v8 Sept 2019)

~20% in 20 sq deg HLV 2019

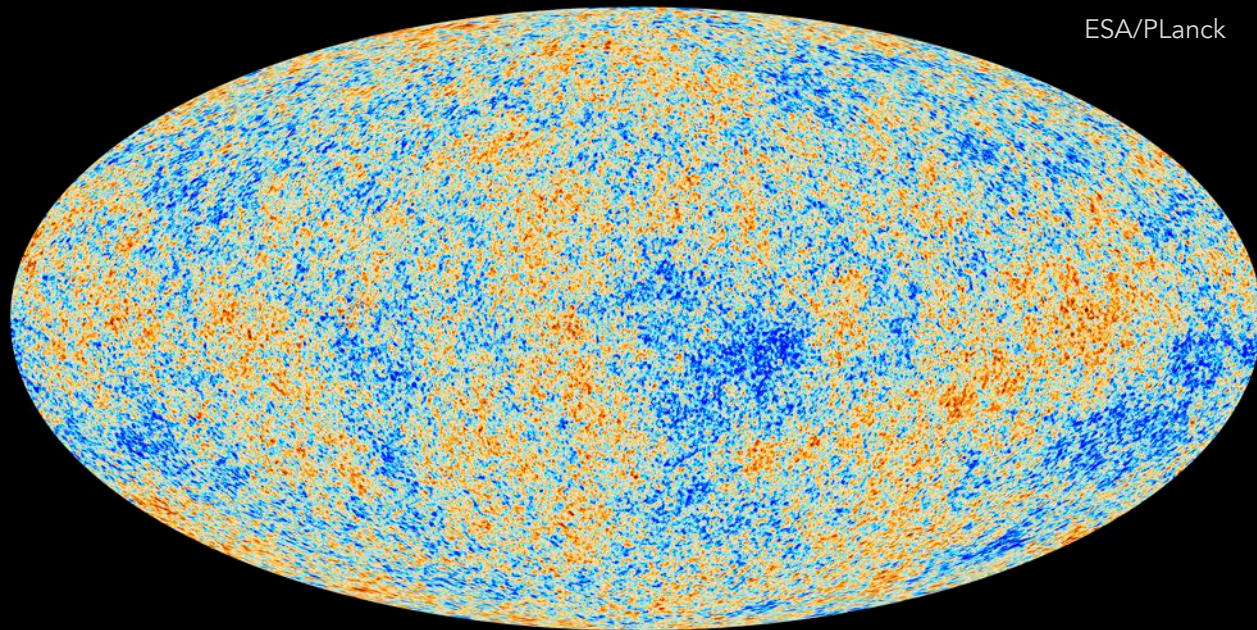


~60% in 10 sq deg HIKLV 2024



Credit: S Fairhurst

ESA/PLanck



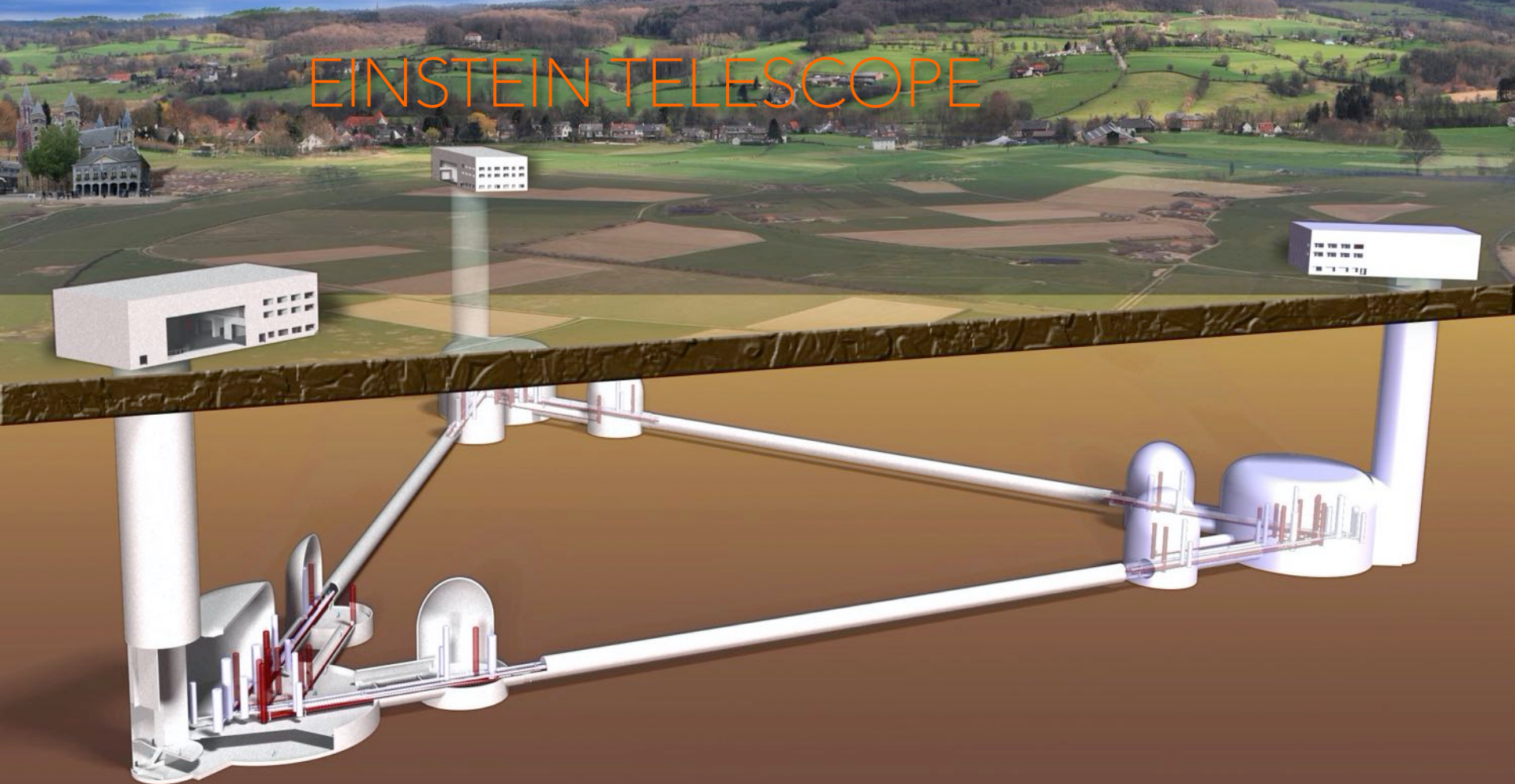
LONG-TERM OUTLOOK

From here to 3G - GWIC

- ❑ 2016: subcommittee to look at the 3G in a coordinated way
 - D Reitze, M Punturo co-chairs
- ❑ 2019: report of actions/recommendations to be delivered
 - Science drivers for 3G detectors
 - Coordination of the ground-based GW community
 - Networking
 - Funding agencies interfacing and advocacy
 - Investigate 3G detectors governance schemes
 - Computing

The GW community is engaged in a coherent effort to develop 3G

EINSTEIN TELESCOPE



EINSTEIN TELESCOPE

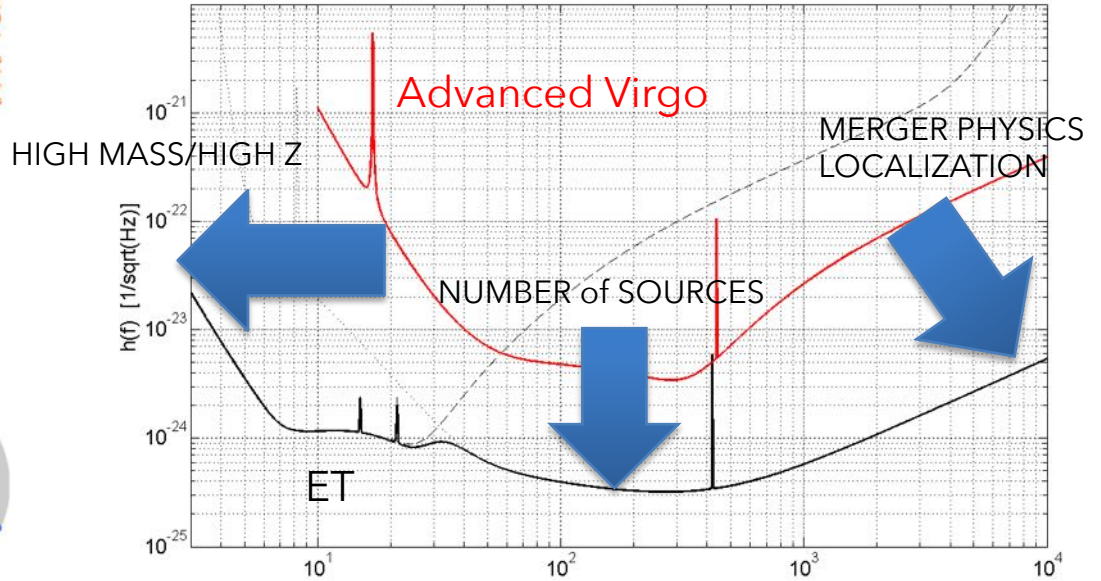
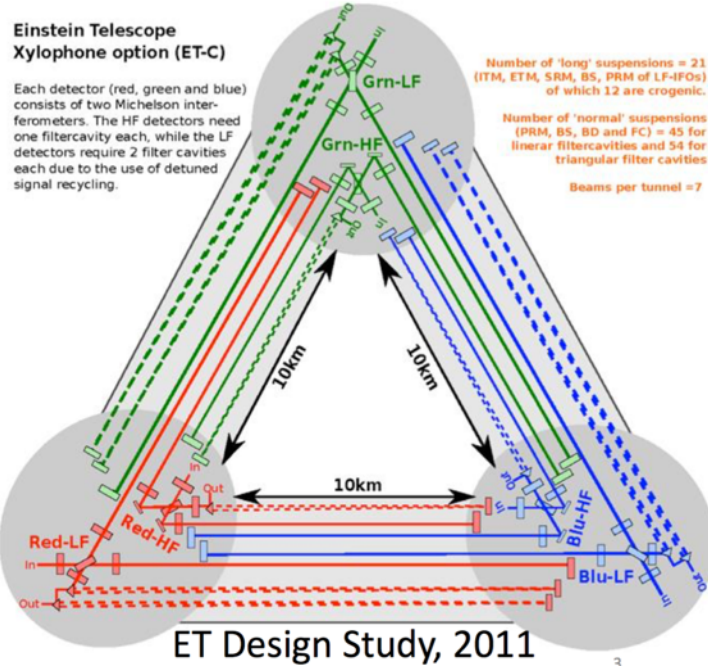
Einstein Telescope Xylophone option (ET-C)

Each detector (red, green and blue) consists of two Michelson interferometers. The HF detectors need one filtercavity each, while the LF detectors require 2 filter cavities each due to the use of detuned signal recycling.

Number of 'long' suspensions = 21
(ITM, ETM, SRM, BS, PRM of LF-IFDs)
of which 12 are cryogenic.

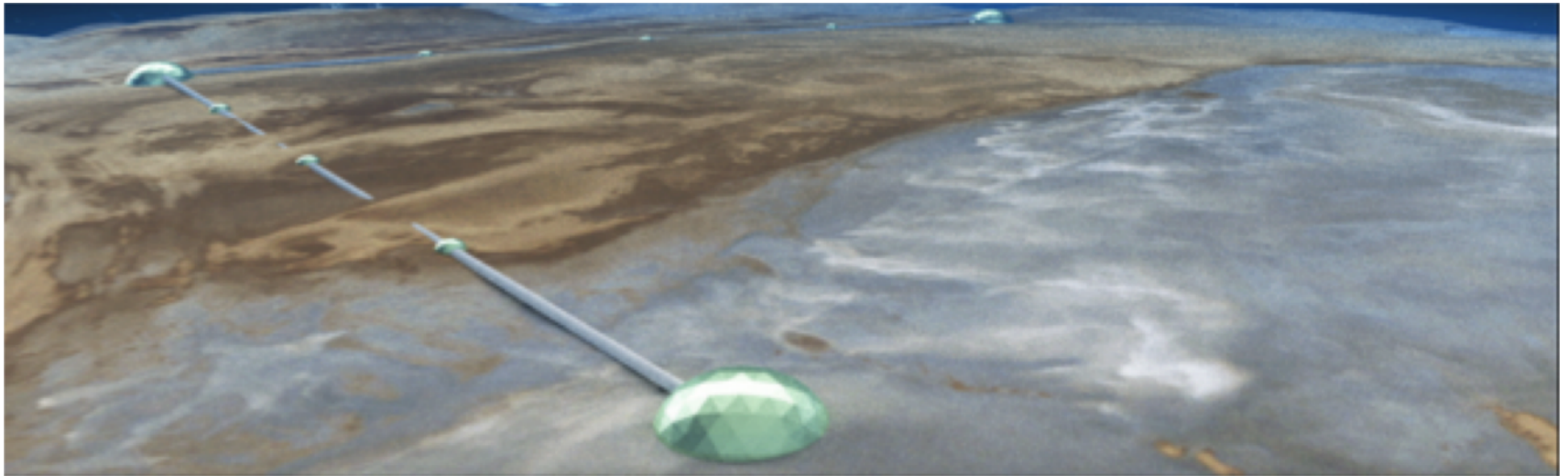
Number of 'normal' suspensions
(PRM, BS, BD and FC) = 45 for
linear filtercavities and 54 for
triangular filter cavities

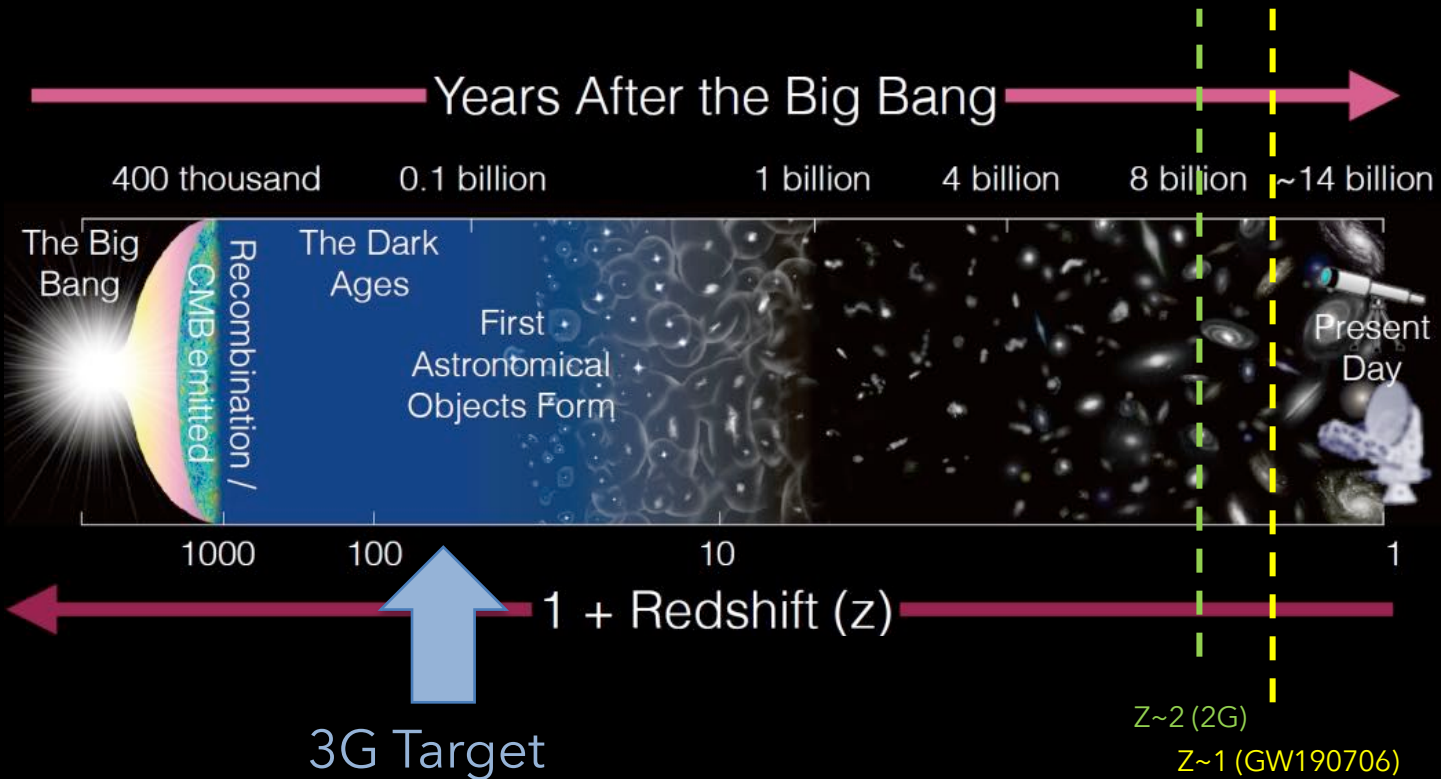
Beams per tunnel = 7



COSMIC EXPLORER

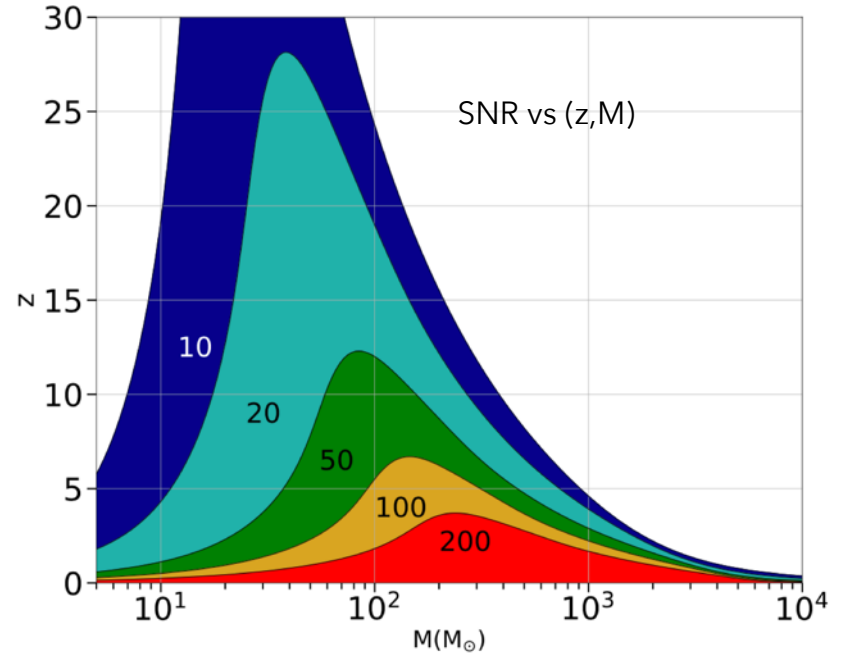
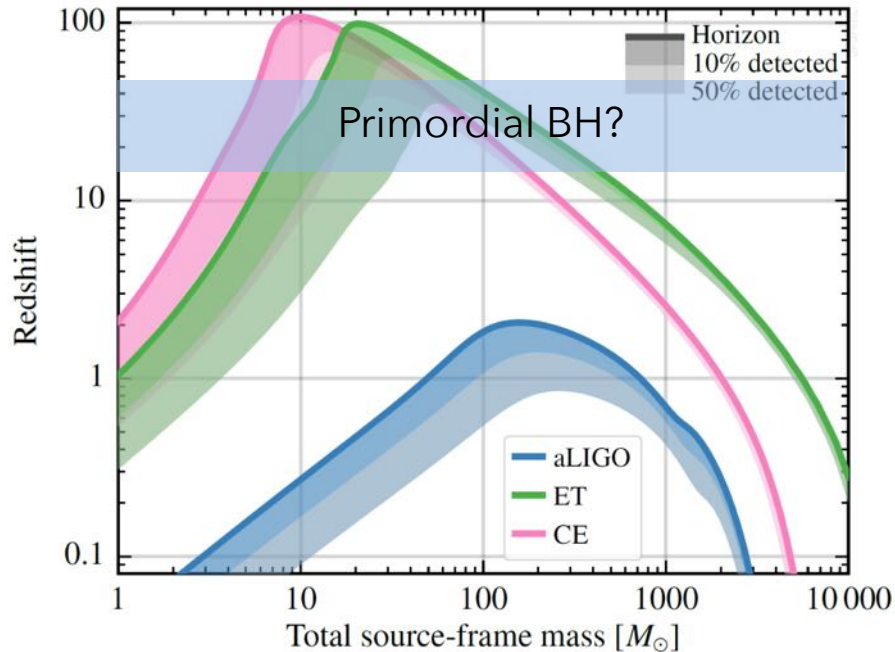
- ❑ 3rd generation observatory in the US
- ❑ Above ground, L shaped, 40 km
- ❑ NSF funded design study under way (www.cosmicexplorer.org)





3G SCIENCE POTENTIAL

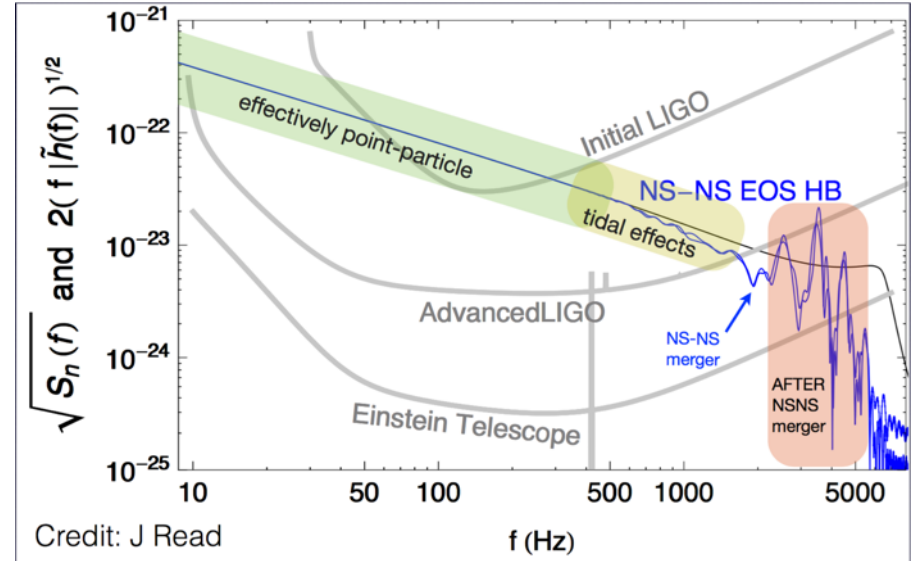
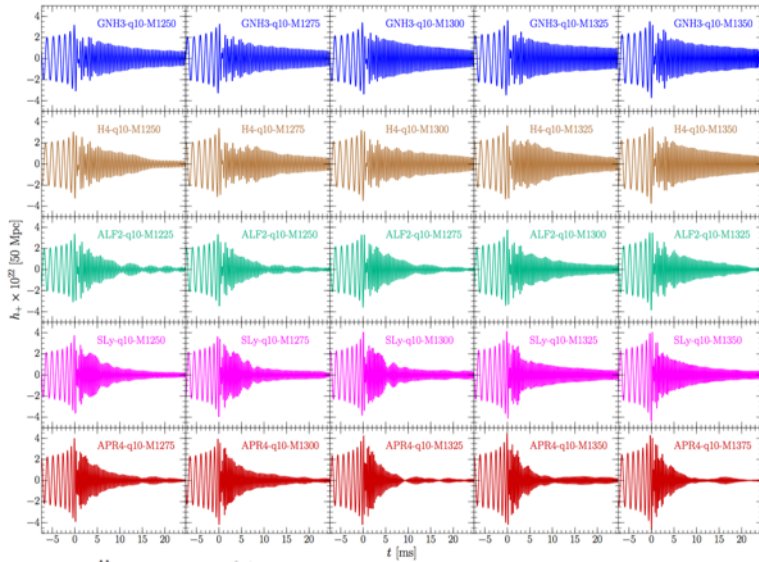
From the ET Science Case



The combination of distances and masses explored, sheer number of detections, and detections with very high SNR will provide a wealth of data that have the potential of triggering revolutions in astrophysics, cosmology and fundamental physics.

NUCLEAR PHYSICS

NS: WE ARE ABLE TO COMPUTE THE WAVEFORMS FOR THE VARIOUS EOS



Credit: J Read

Takami, Rezzolla, Baiotti (2014)

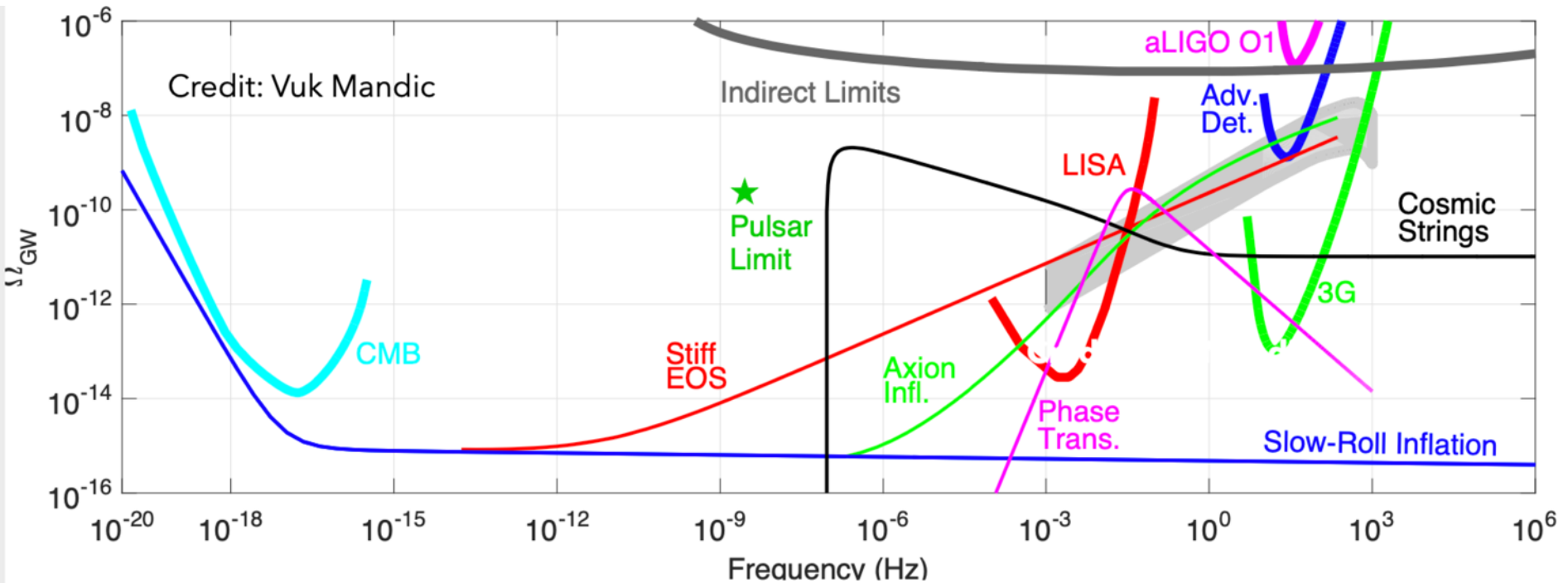
A 3G DETECTOR IS NEEDED TO MEASURE WHICH EOS IS THE RIGHT ONE

COSMOLOGY

$$d_L(z) = \frac{1+z}{H_0} \int_0^z \frac{dz'}{\sqrt{\Omega_M(1+z')^3 + \frac{\rho_{DE}(z')}{\rho_0}}}$$

- ❑ GW170817 has already demonstrated the possibility to measure H_0 with GW standard sirens
- ❑ Multi-messenger observations at high redshift with 3G will make possible to measure $\rho_{DE}(z)$
- ❑ Possibility to explore models beyond Λ CDM

COSMOLOGY

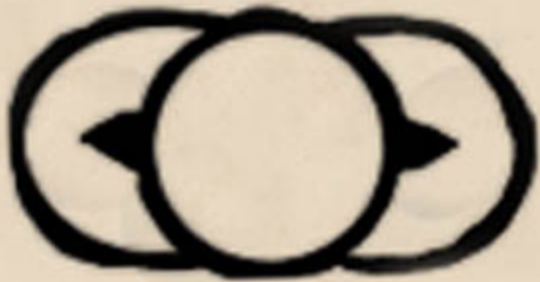




Galileo, 1610



Galileo, 1610

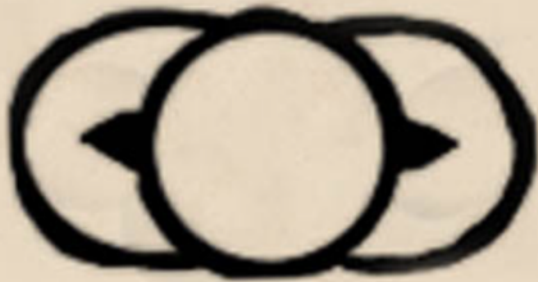


Galileo, 1616

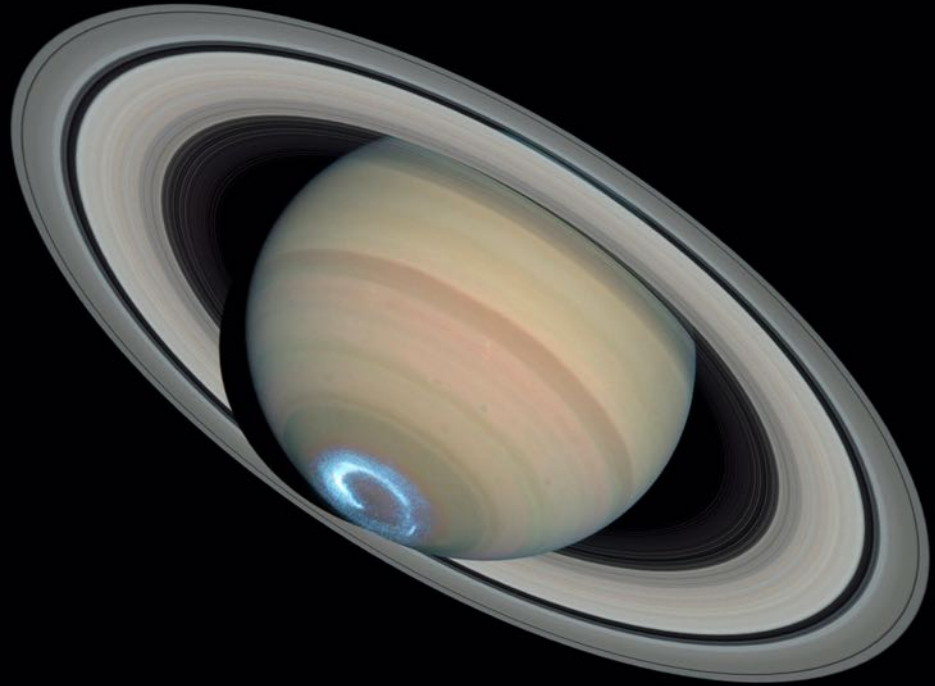
WE HAVE THE RIGHT INSTRUMENT.
NOW WE NEED TO MAKE IT BETTER AND BETTER AND BETTER...



Galileo, 1610

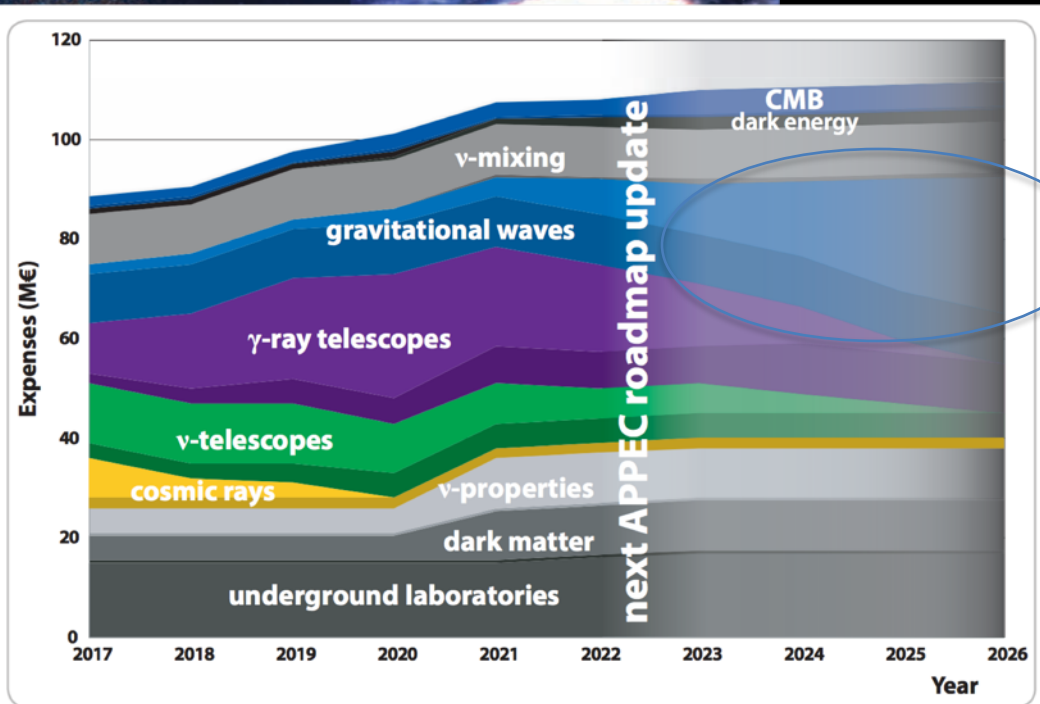


Galileo, 1616



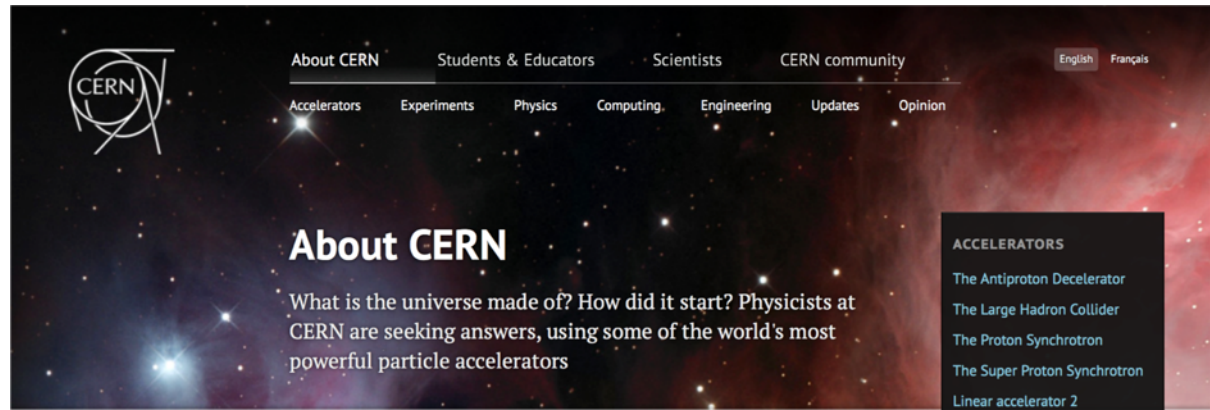
HST, 400 yrs later

European Astroparticle Physics Strategy 2017-2026



THE POSSIBLE ROLE OF CERN

- ❑ The GW community looks at CERN as a model to many extents
- ❑ We have a lot to learn from CERN:
 - Model of governance
 - Management of big projects
 - Technology: underground infrastructure, vacuum, cryogenics, computing
- ❑ A commitment of CERN on ET (in some form) might be a game changer



From the Physics Briefing Book

<https://home.cern/news/news/cern/briefing-book-2020-update-european-strategy-particle-physics>

The future generation of GW ground-based detectors, which in Europe is the Einstein Telescope project, has the unique potential **to strengthen the synergy between cosmology, particle, nuclear, astroparticle physics and astrophysics science domains** since: 1) It **tests the nature of gravity** (...) 2) It **explores the nature of dark matter** in a complementary way to colliders and underground direct search experiments (...); 3) It **provides an independent measure of the Hubble constant** either solving the current tension between its *far* and *near* determinations or alternatively opening a portal to new physics (...); 4) It reveals phase transition from nucleons to free quarks giving **insight into the QCD phase diagram**, explores the state of ultra-dense nucleons and the origin of heavy elements (...); 5) It **studies the primordial Universe**, through primordial stochastic backgrounds, early-Universe phase transitions, cosmic strings (...); 6) It **probes the nature of space-time at the interface with quantum mechanics**, through the study of alternative gravity and quantum gravity theories.

ET science not only reflects CERN's own pursuit of fundamental physics, but strong synergies exist concerning the precision engineering that was required to build and operate the LHC.

CONCLUDING REMARKS

- ❑ A new field of research has been opened promising access to fundamental physics, nuclear physics, astrophysics, cosmology, dark sector, ...
- ❑ Time to invest on the development of 3G technologies, infrastructure, detectors
- ❑ Einstein Telescope: a EU project for a 3G detector of GW
 - Very strong science case
 - Cost ~ 1.5 G€
 - Timeline > 2030
- ❑ A great chance to ensure EU leadership in the field